

# Inclusive Overview



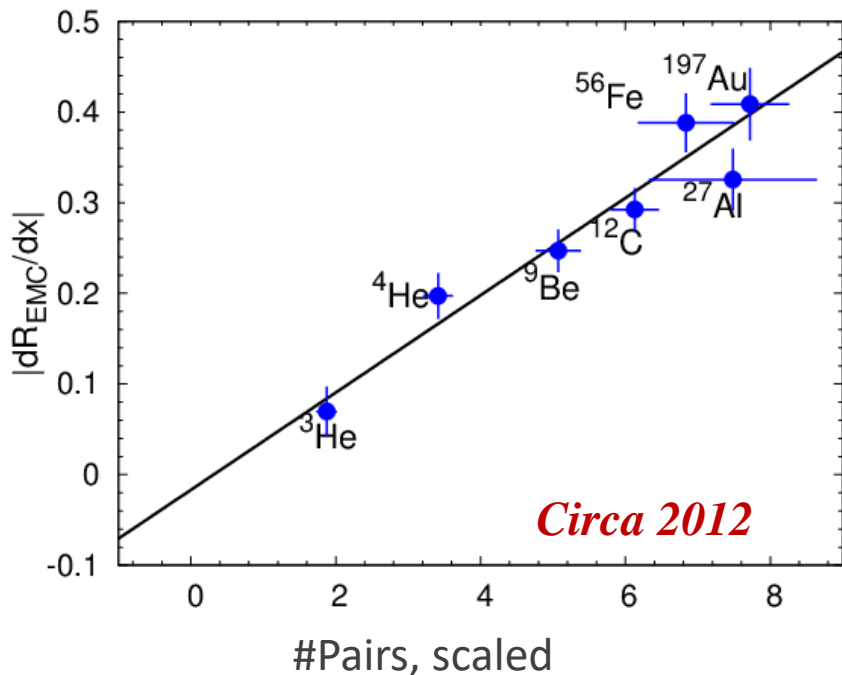
3/20/2019

*2<sup>nd</sup> Workshop on Quantitative Challenges in SRC and EMC research*

**Nadia Fomin**



# Inclusive data is the source of a lot of info on SRC/EMC



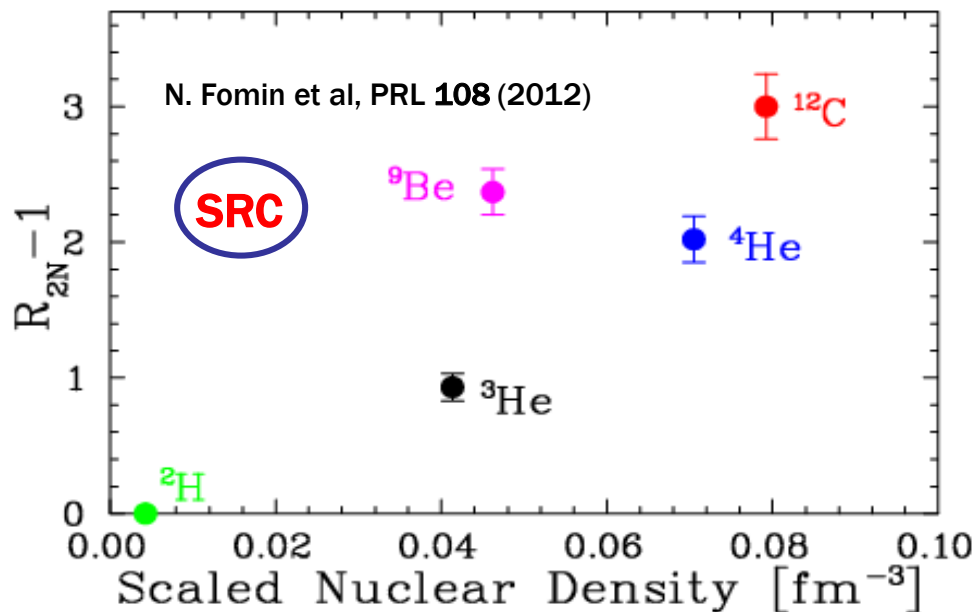
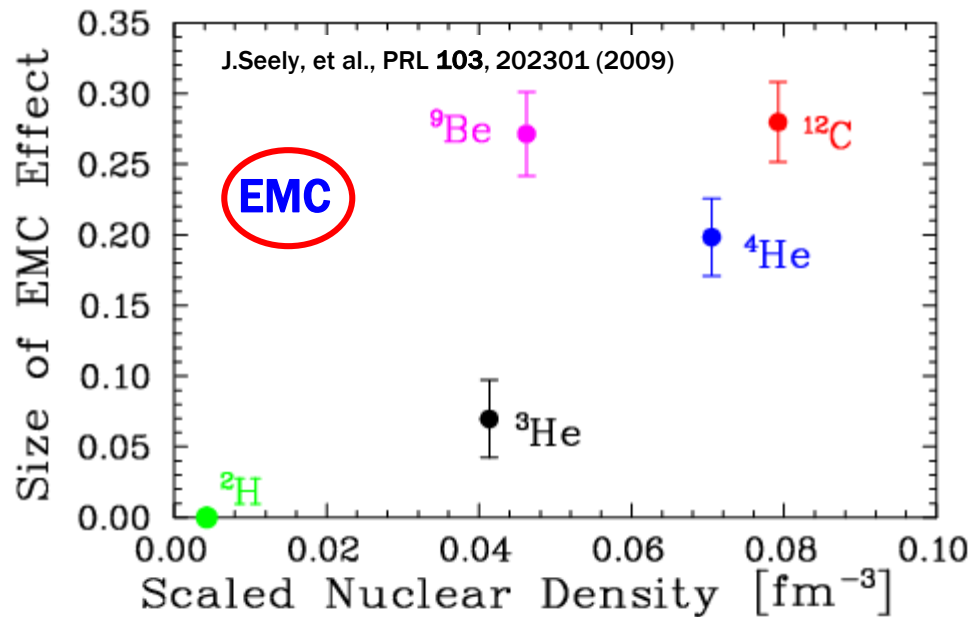
J. Seely, et al., PRL103, 202301 (2009)

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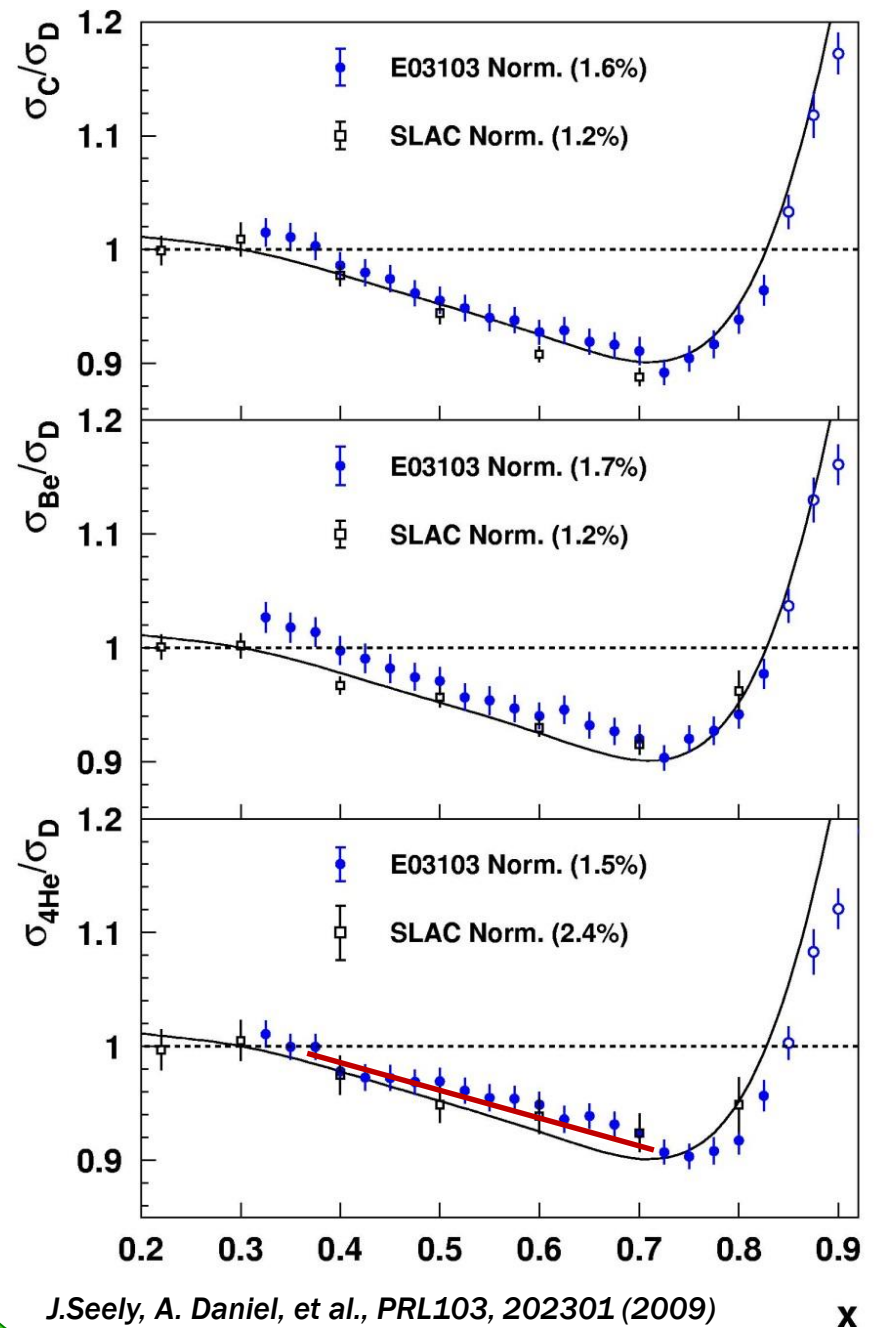
# Precision results on light nuclei from JLab E03-103

- $C/D$  and  ${}^4\text{He}/D$  ratios – no isoscalar correction necessary
- Consistent with SLAC results, but much higher precision at high  $x$

- Fit the slope of the ratios for  $0.35 < x < 0.7$ :

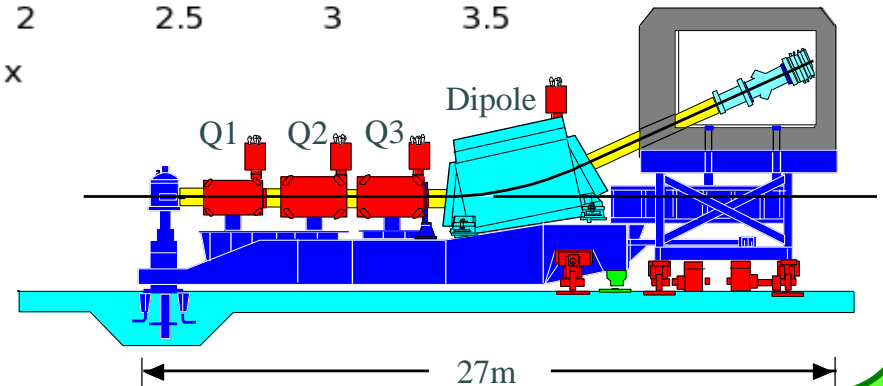
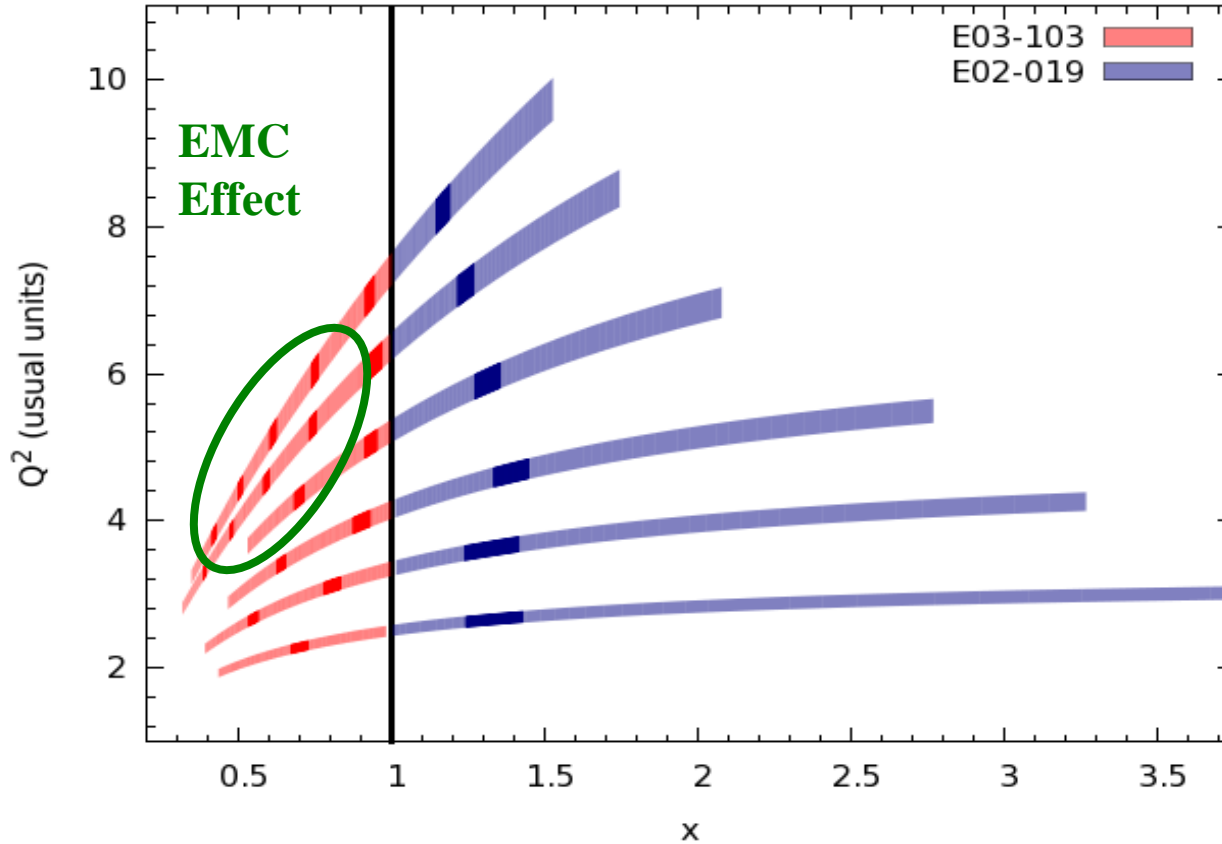
$$\frac{dR_{EMC}}{dx}$$

- Compare across nuclei

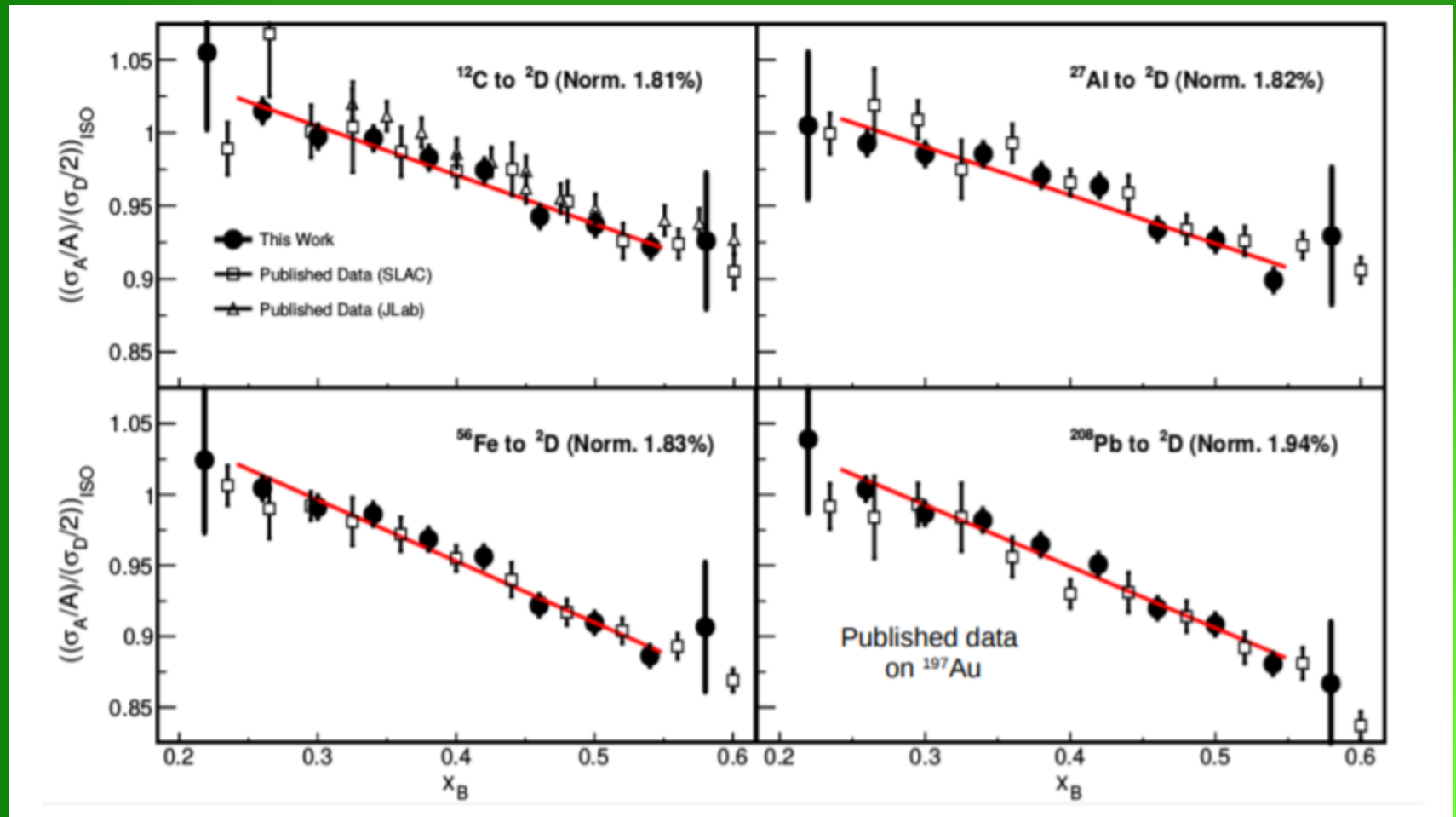


# Hall C, high-precision inclusive program at 6GeV

$^2\text{H}$   
 $^3\text{He}$   
 $^4\text{He}$   
 $^9\text{Be}$   
 $^{12}\text{C}$   
 $^{27}\text{Al}^*$   
 $^{63}\text{Cu}$   
 $^{197}\text{Au}$

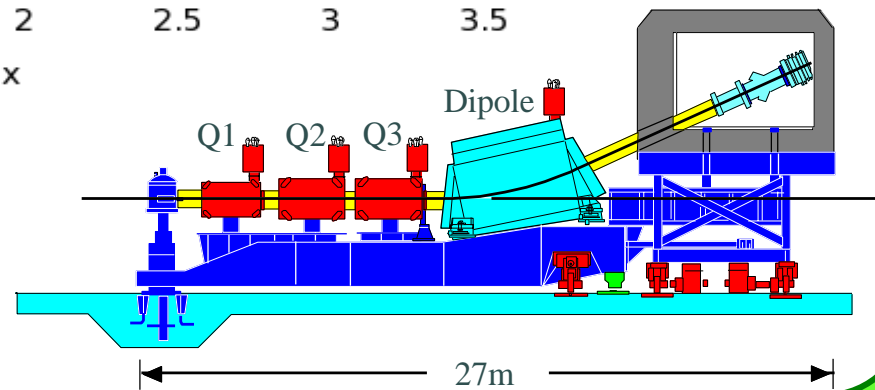
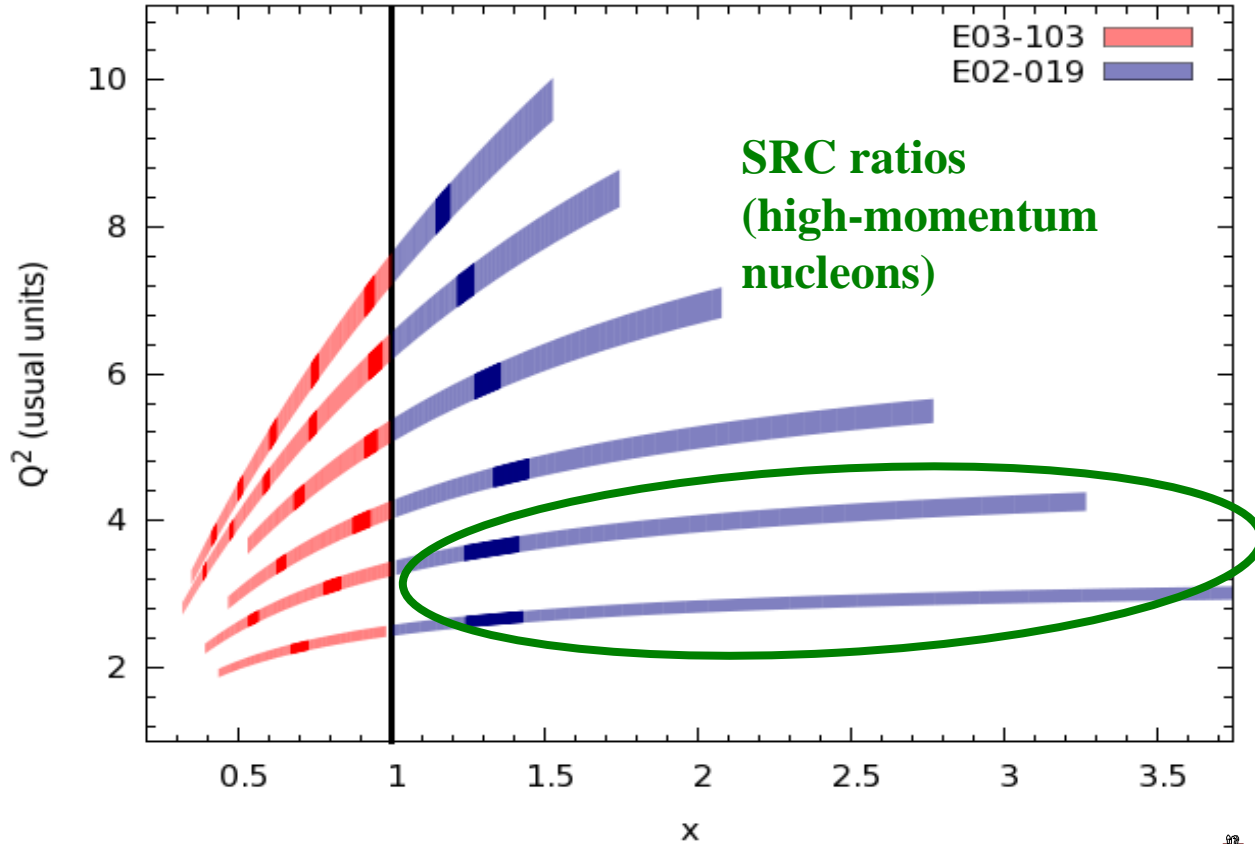


# New EMC Ratios from CLAS data mining project



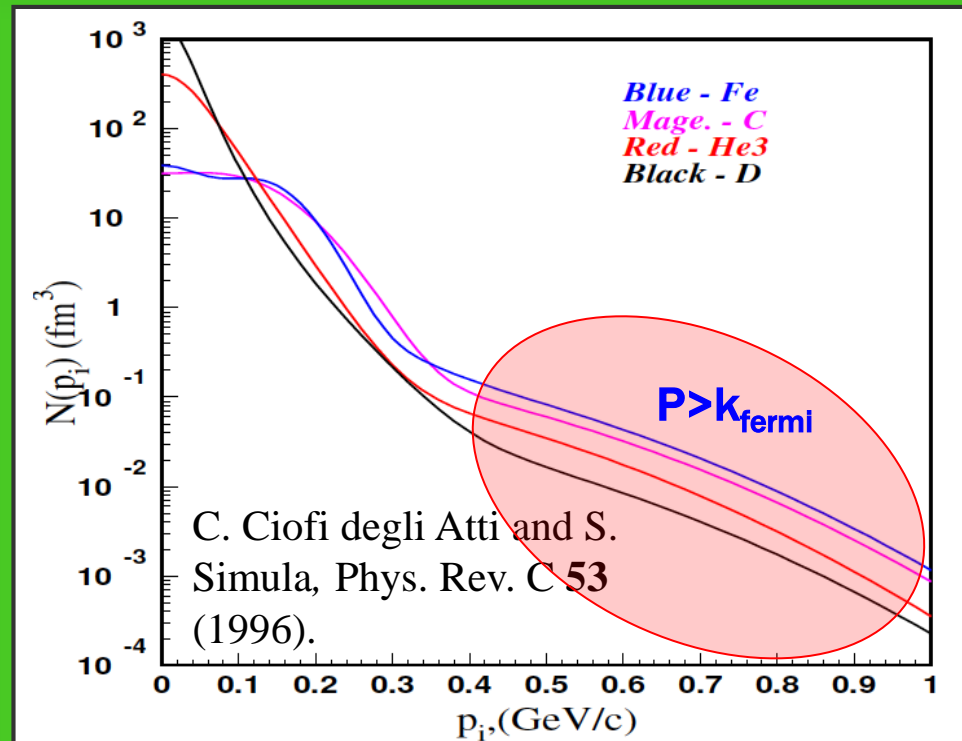
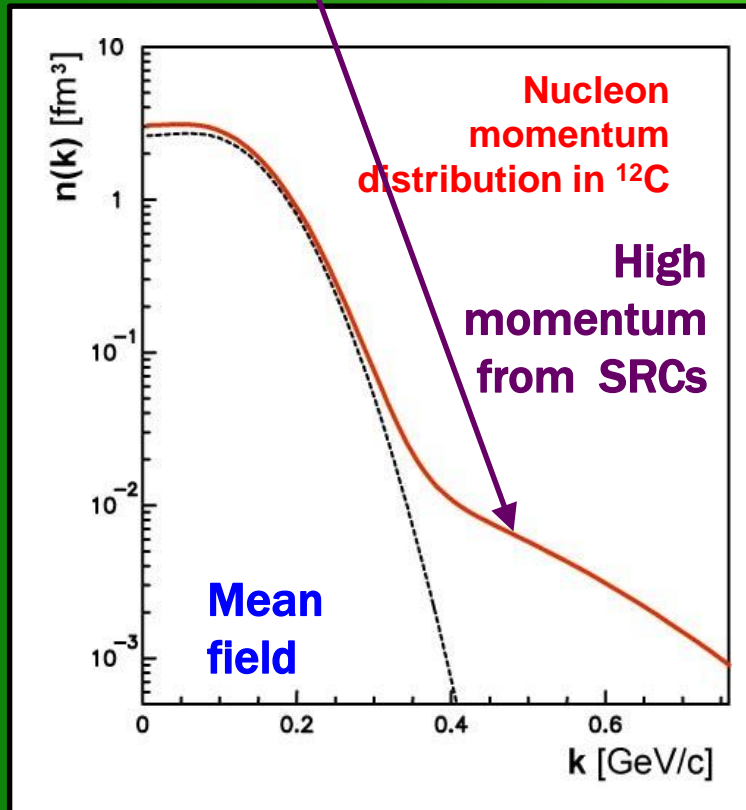
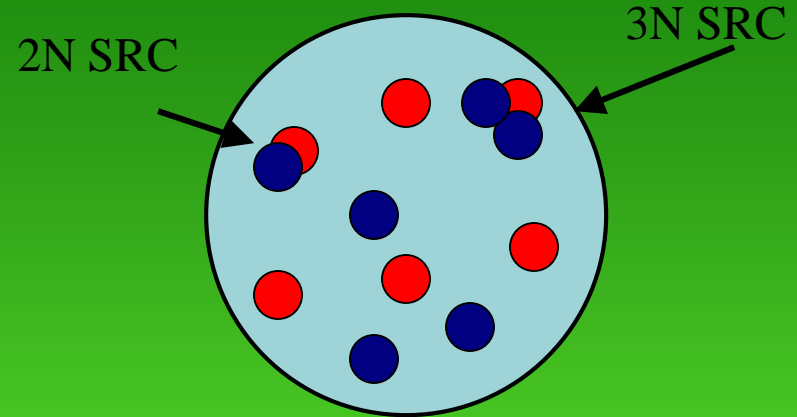
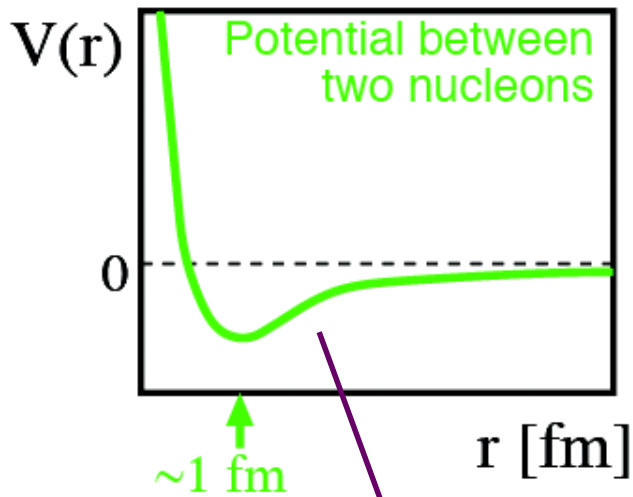
# Hall C, high-precision inclusive program at 6GeV

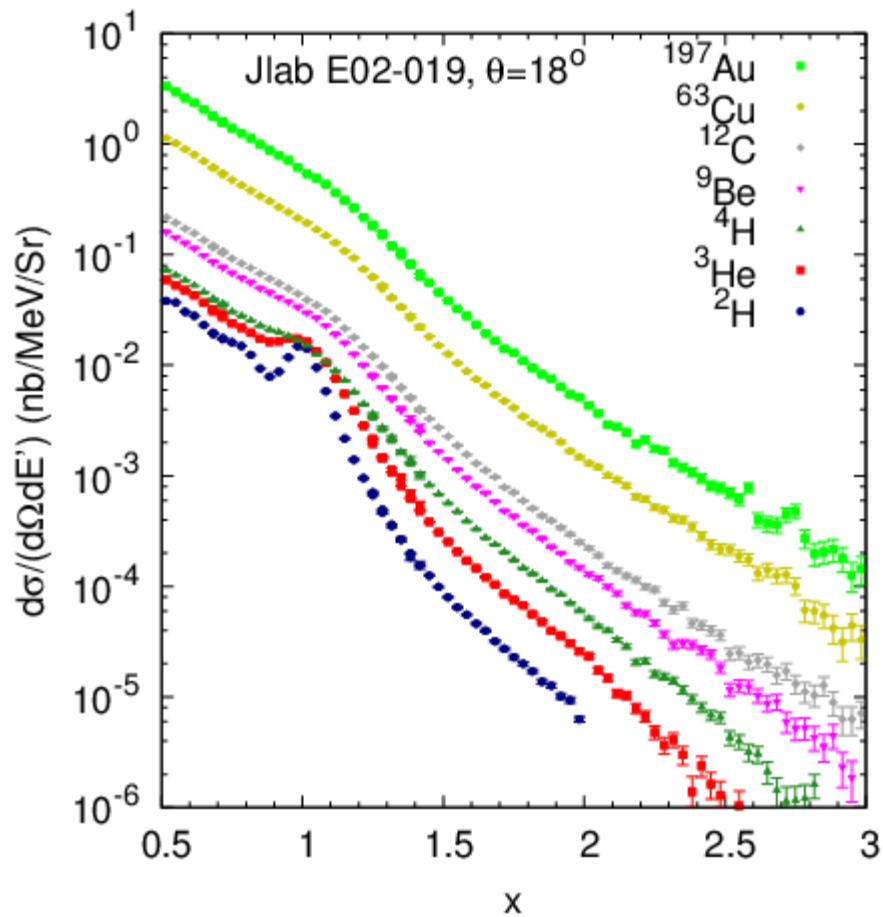
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# High momentum nucleons

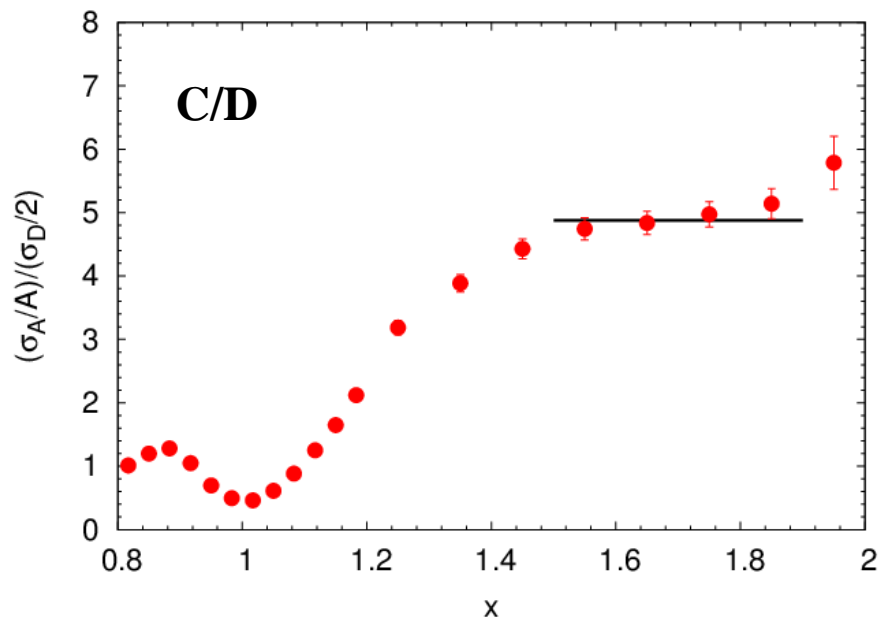
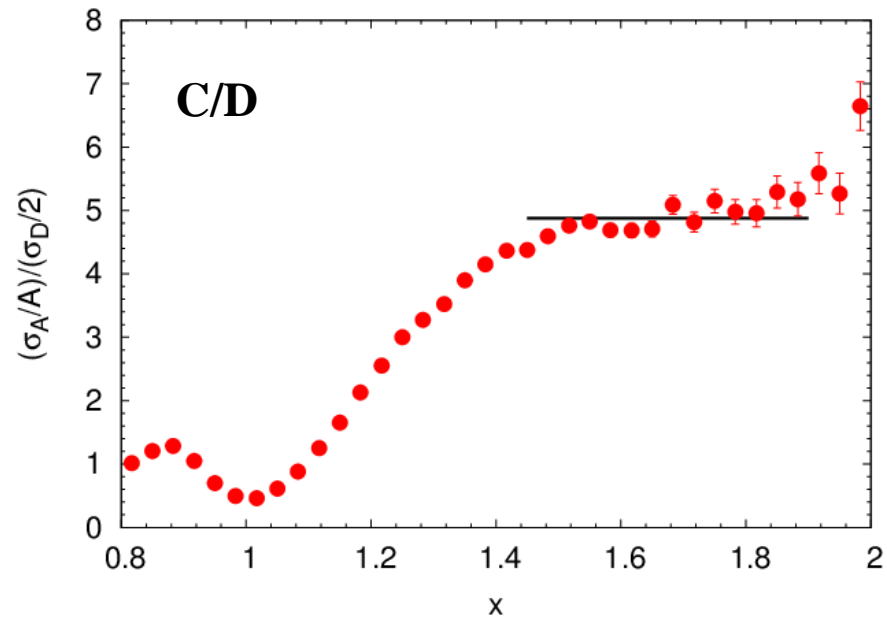
## - Short Range Correlations





Cross section binning:  $x_{\text{binsize}}=0.033$

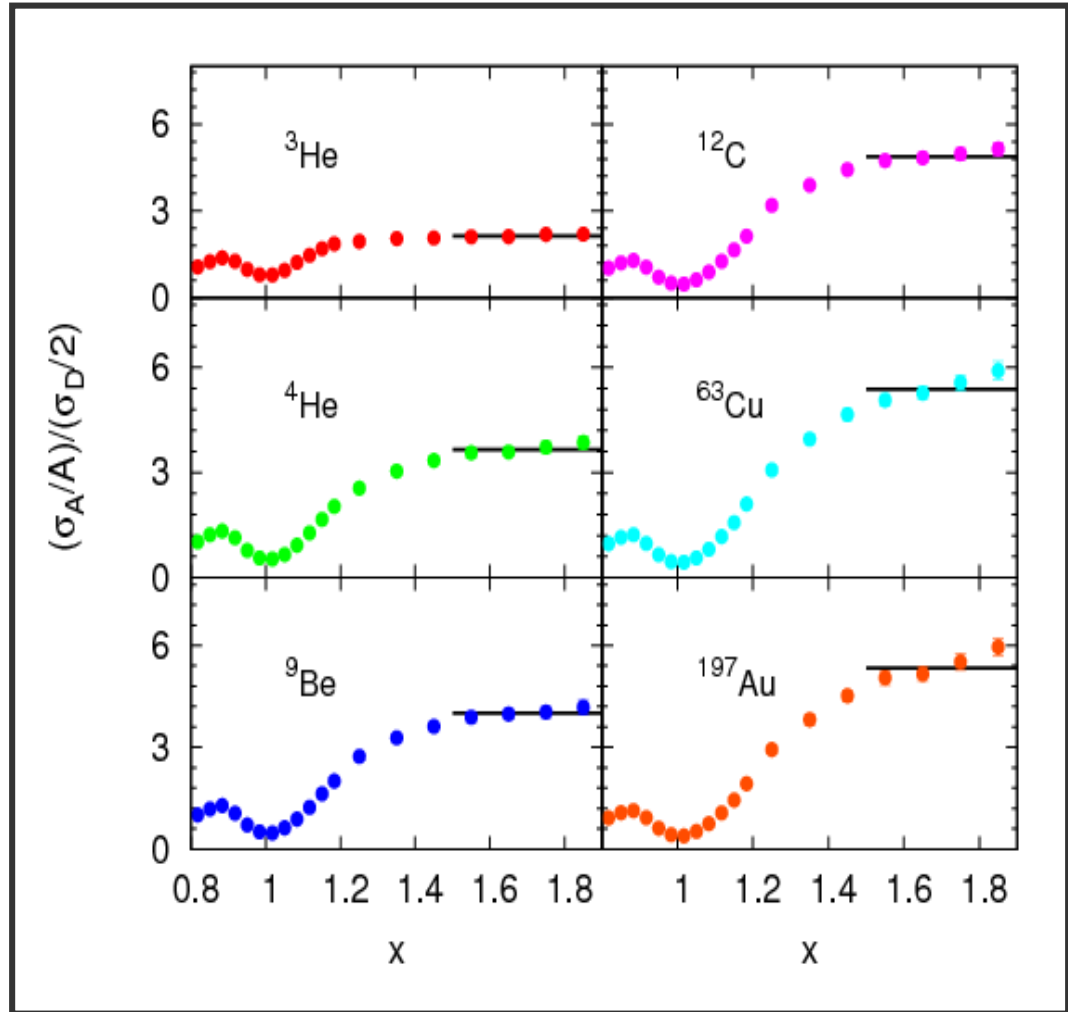
Ratios rebinned starting at  $x=1.2$





# E02-019: 2N correlations in A/D ratios

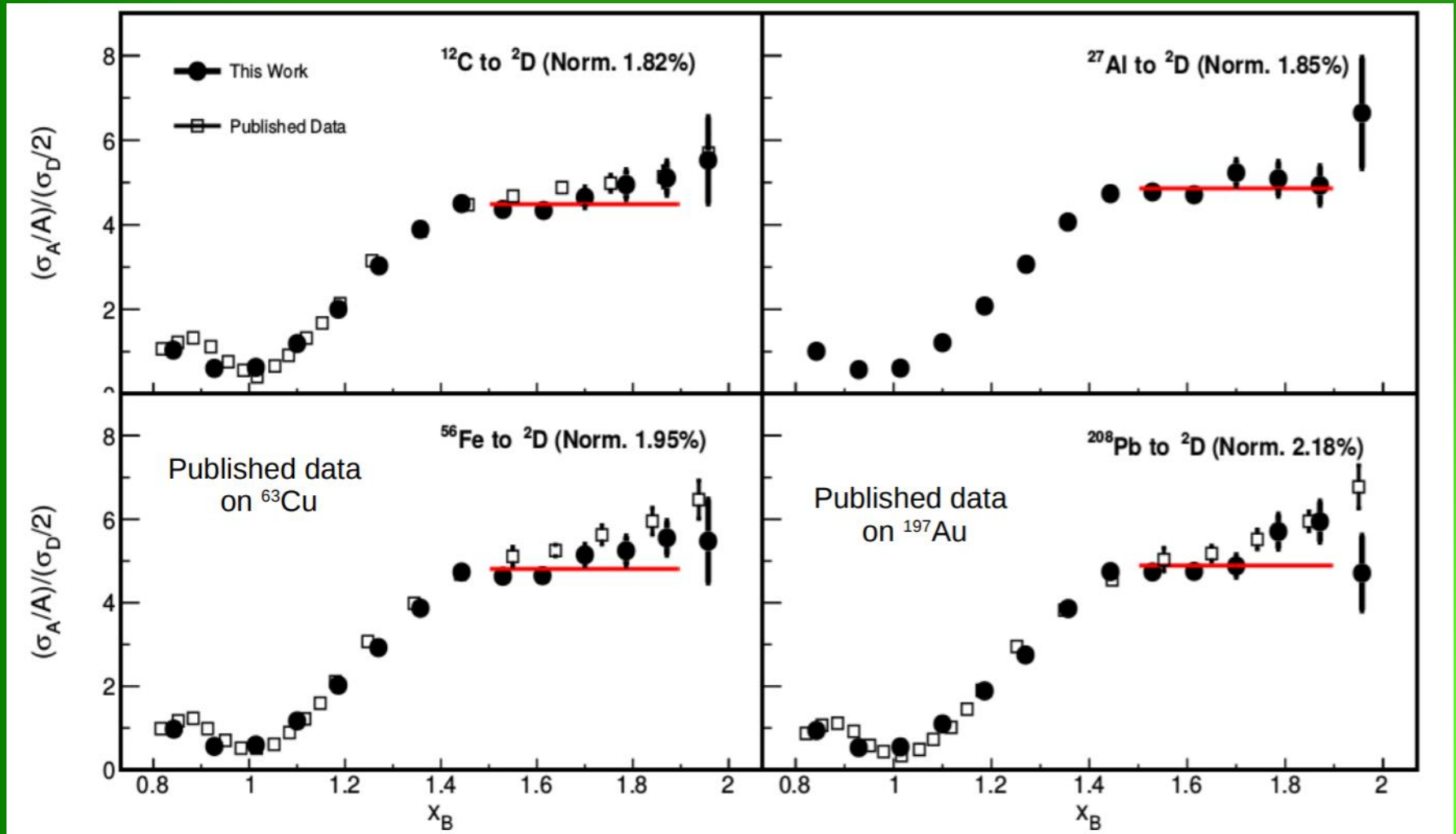
A	$\theta_e = 18^\circ$
$^3\text{He}$	$2.14 \pm 0.04$
$^4\text{He}$	$3.66 \pm 0.07$
Be	$4.00 \pm 0.08$
C	$4.88 \pm 0.10$
Cu	$5.37 \pm 0.11$
Au	$5.34 \pm 0.11$
$\langle Q^2 \rangle$	$2.7 \text{ GeV}^2$
$x_{\text{min}}$	1.5



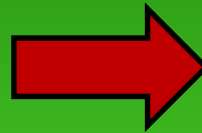
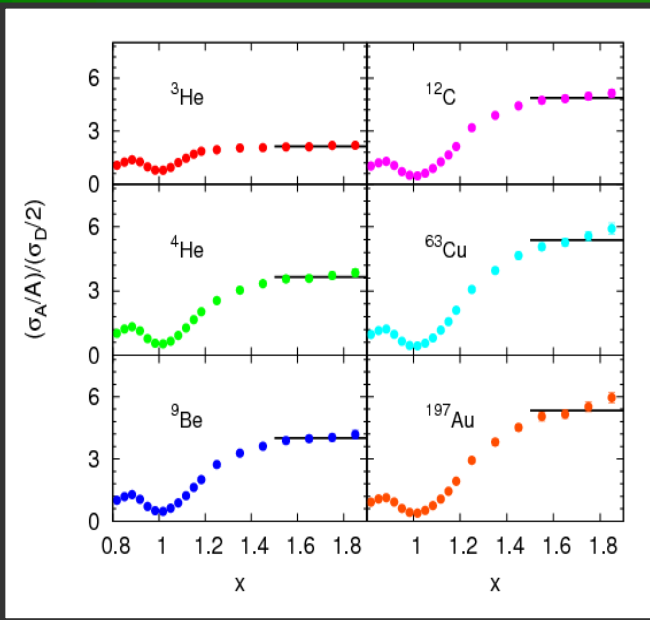
- 5% of strength in high momentum tail (calculated)
- 20% of nucleons spend time in SRC configurations

$$\langle Q^2 \rangle = 2.7 \text{ GeV}^2$$

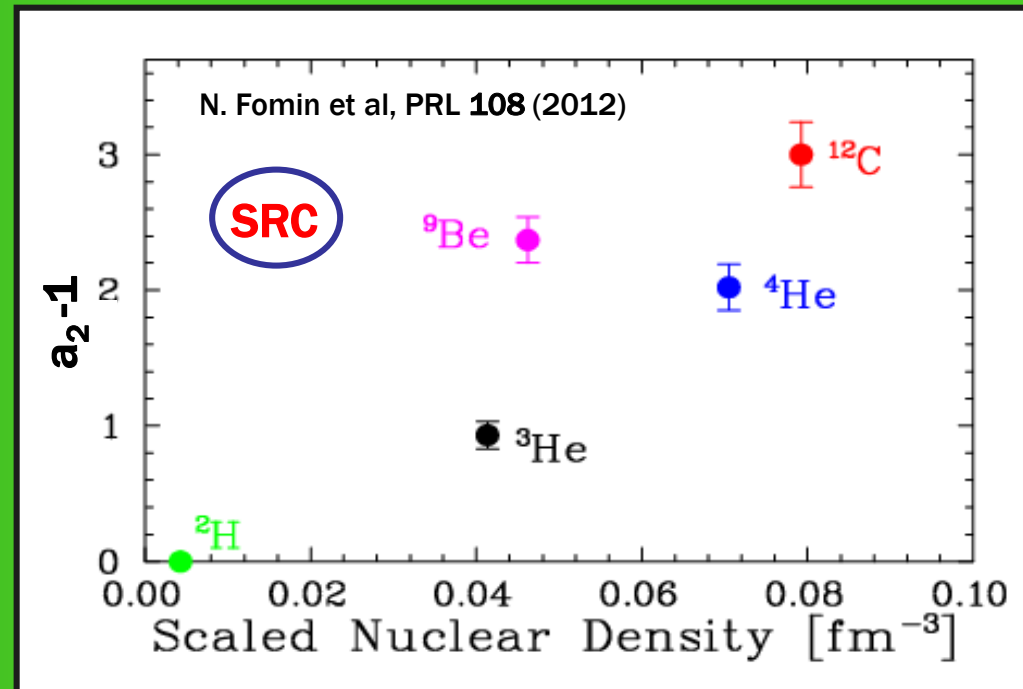
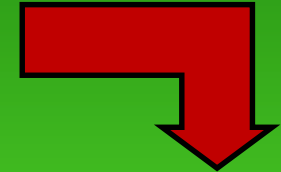
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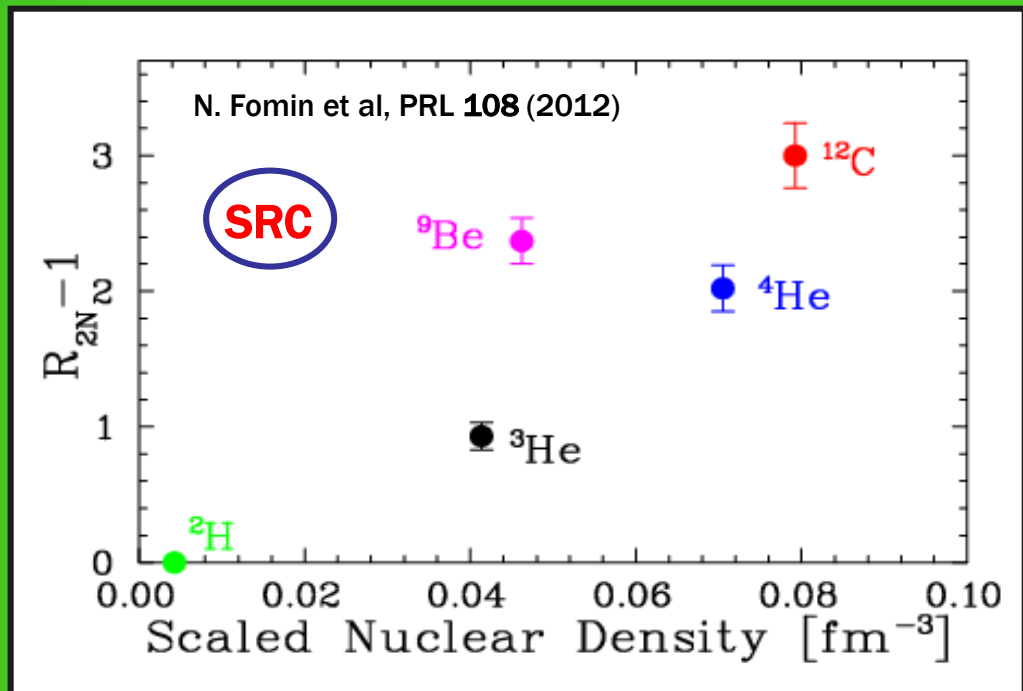
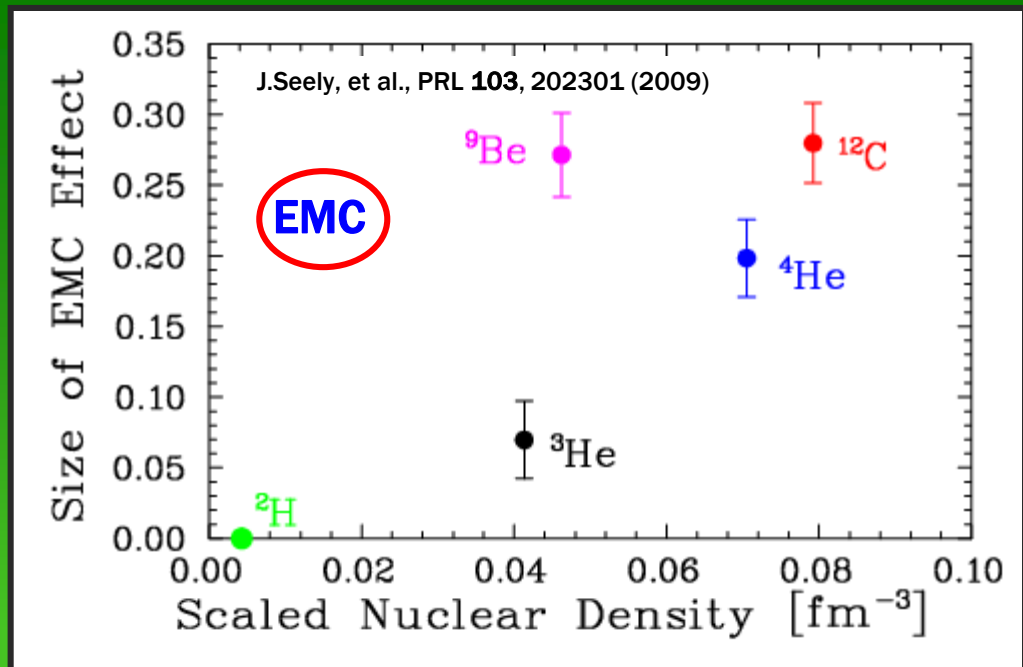


# Look at nuclear dependence of NN SRCs

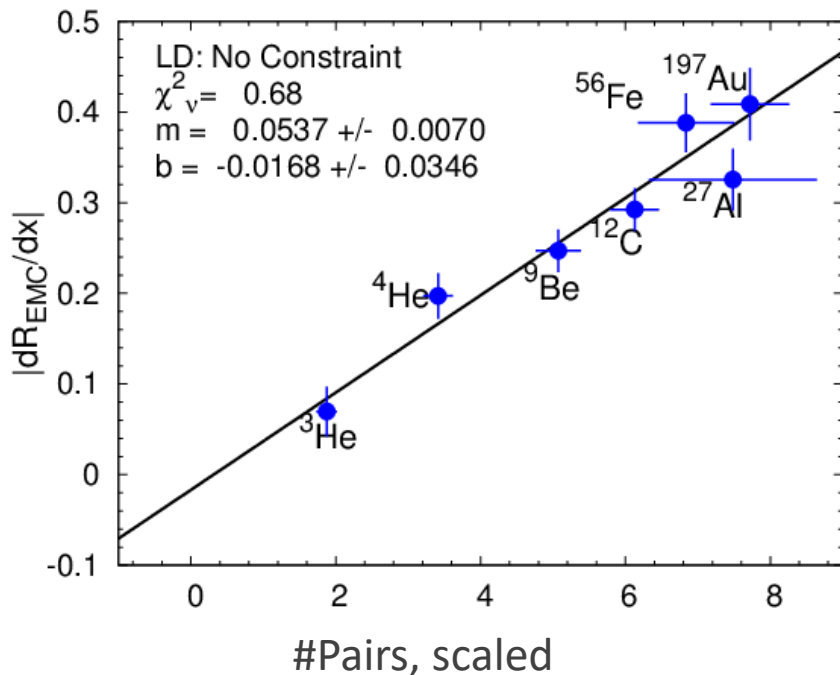


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# Thanks, ${}^9\text{Be}$ !



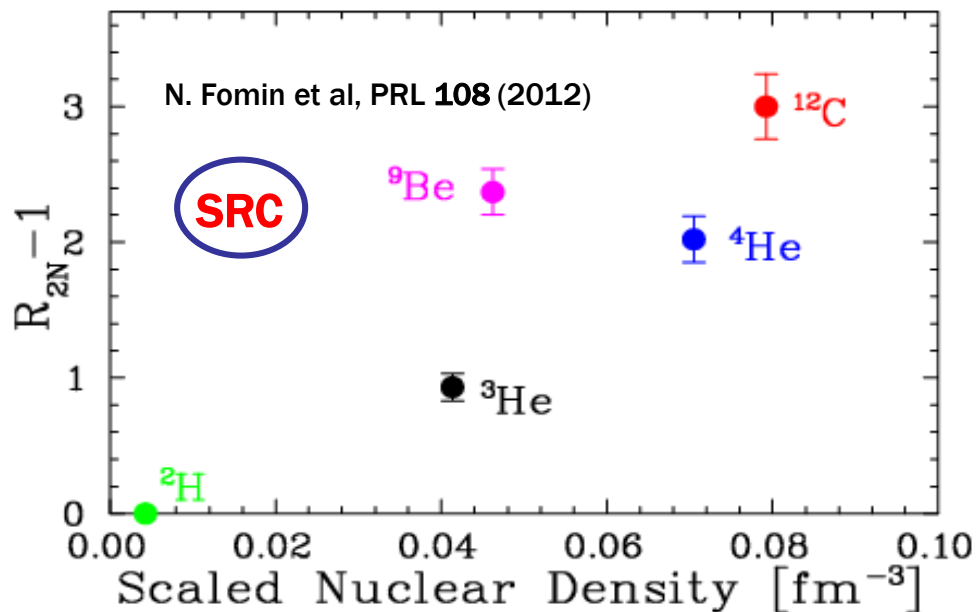
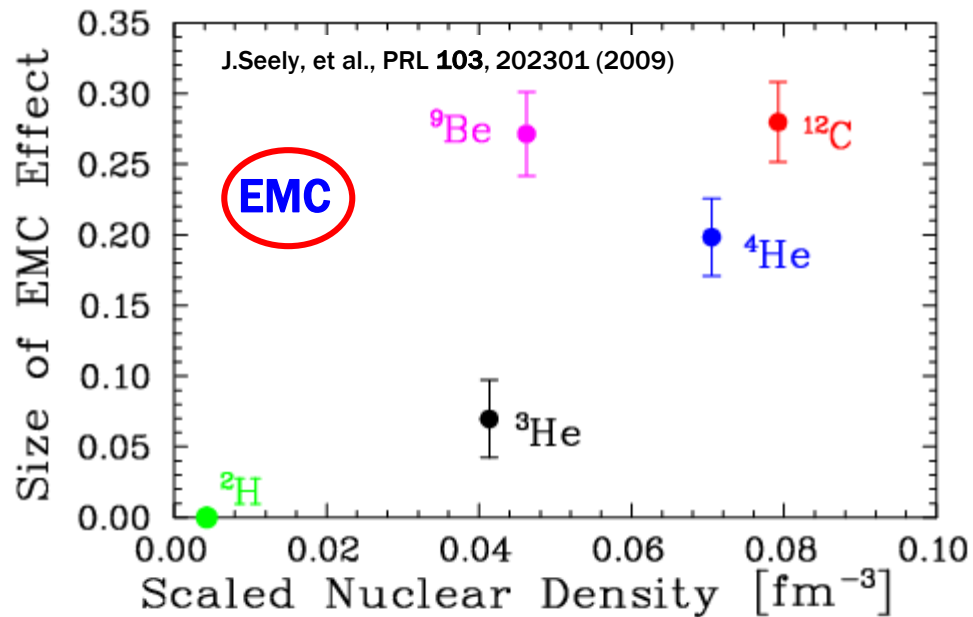
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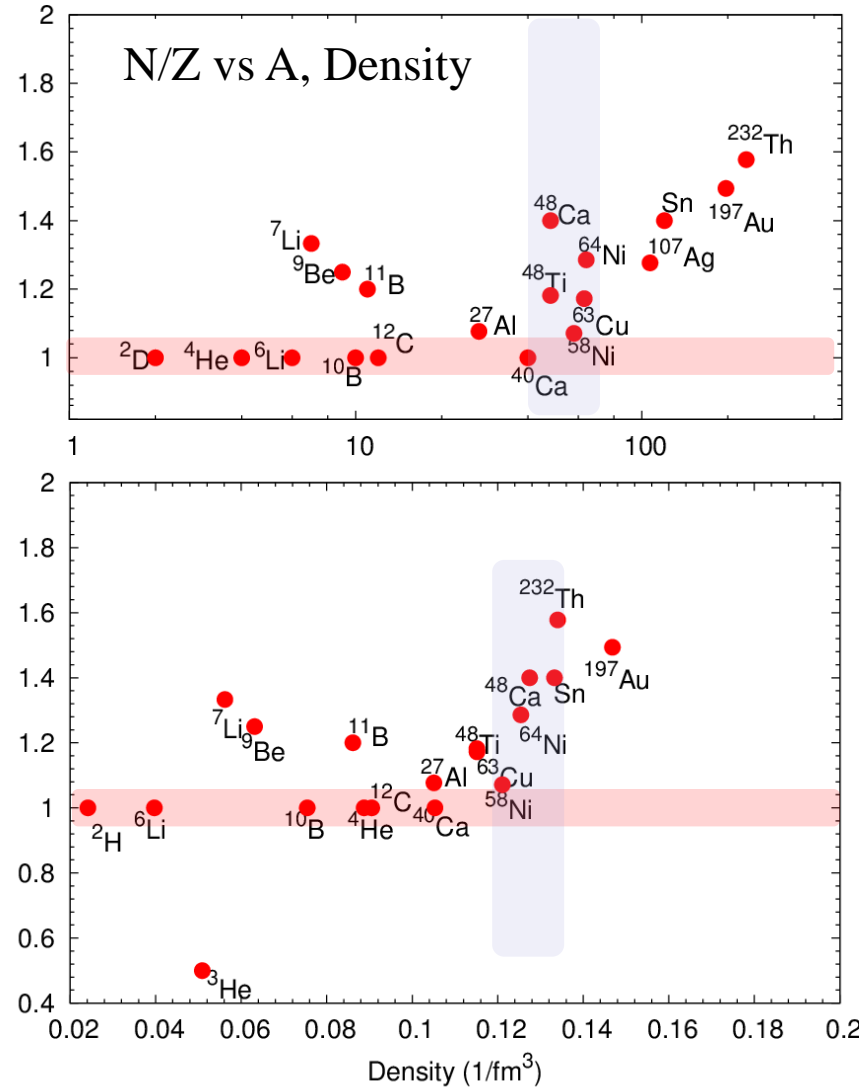
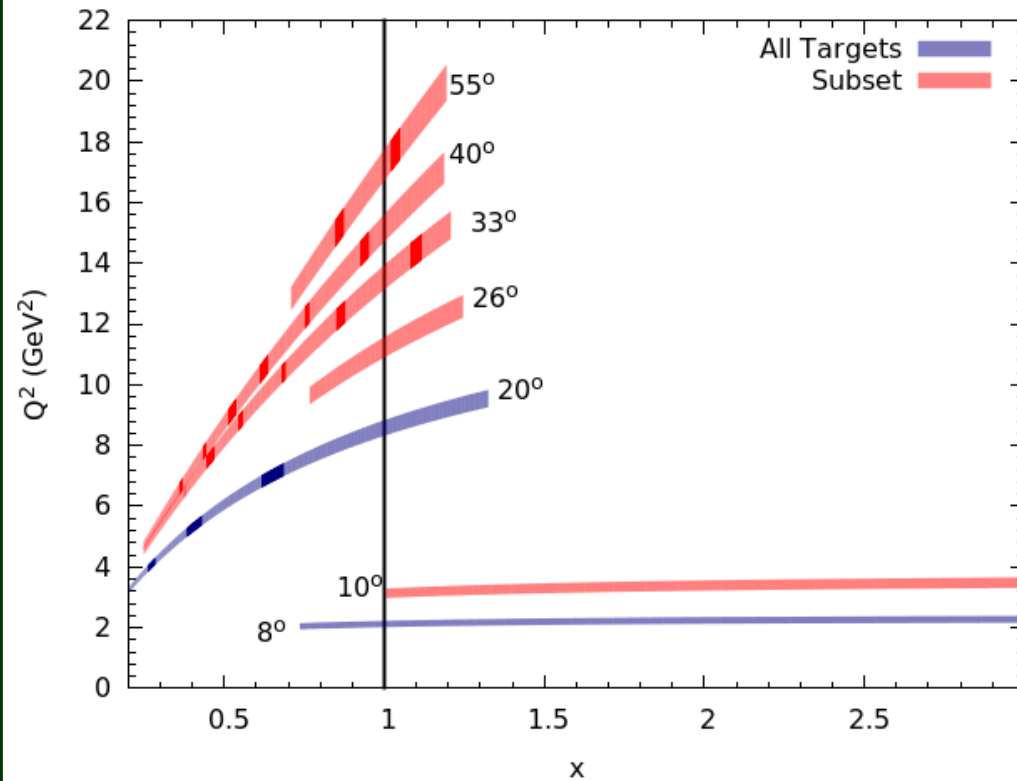
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L. Weinstein, et al., PRL 106, 052301 (2011)



# 12 GeV: The next inclusive generation

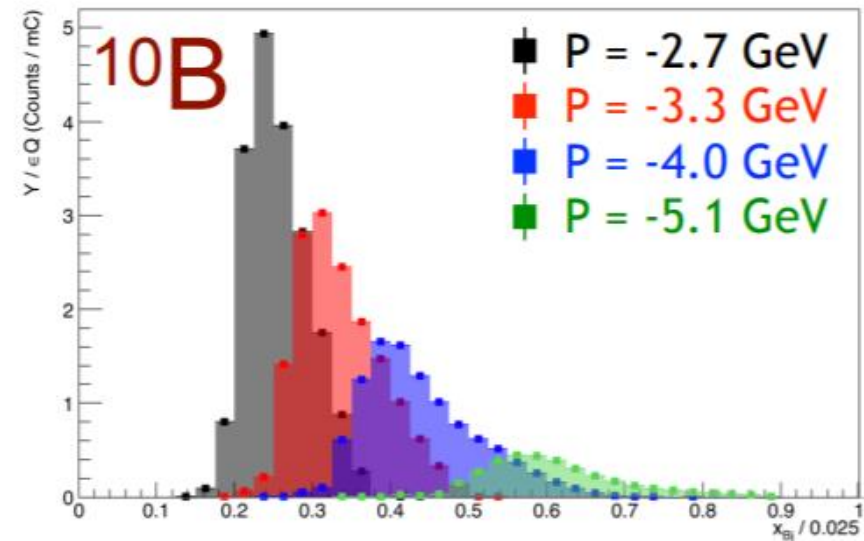
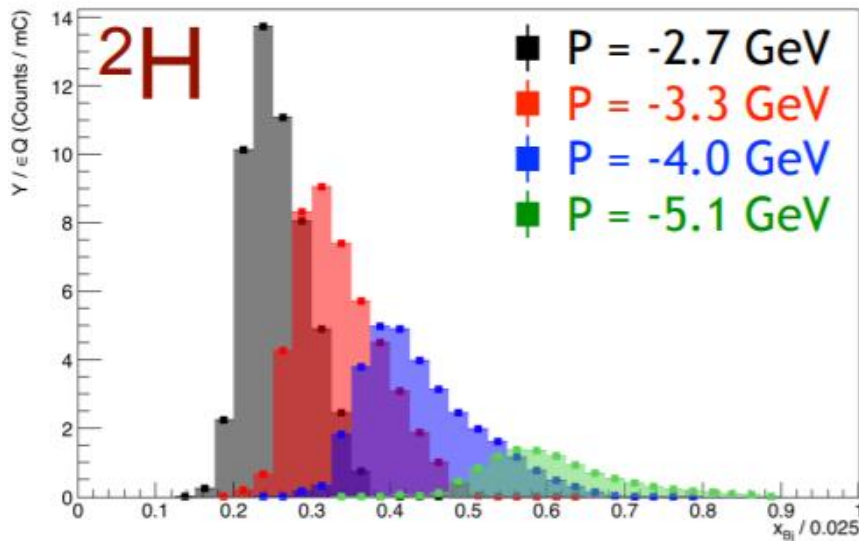
## Target Choice motivated by physics impact



# Hall C data on new nuclei!

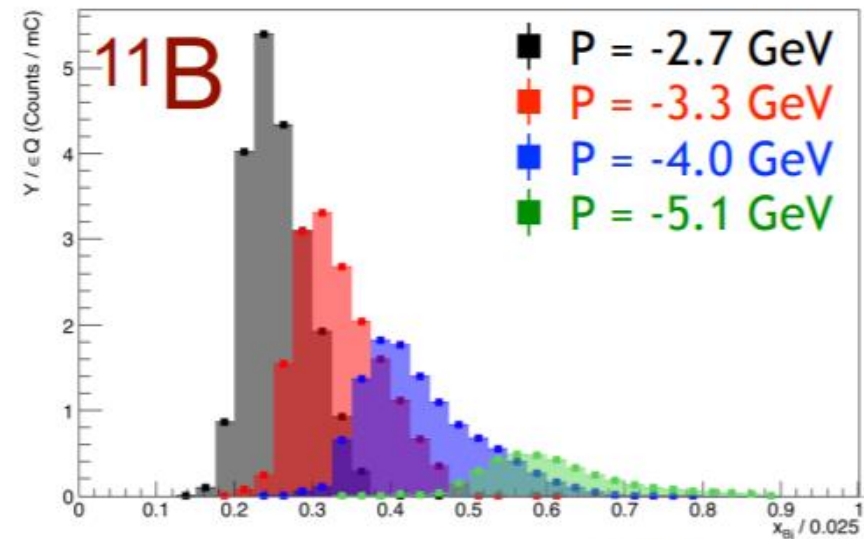


# Hall C data on new nuclei!



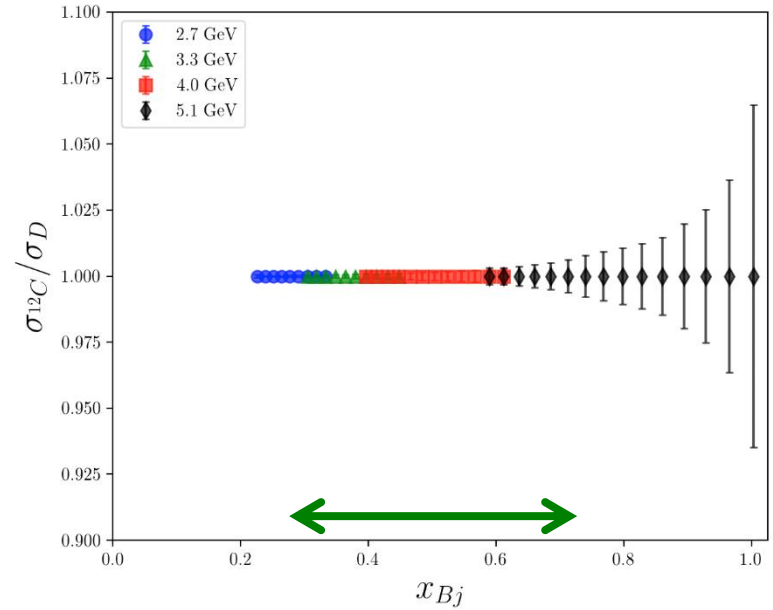
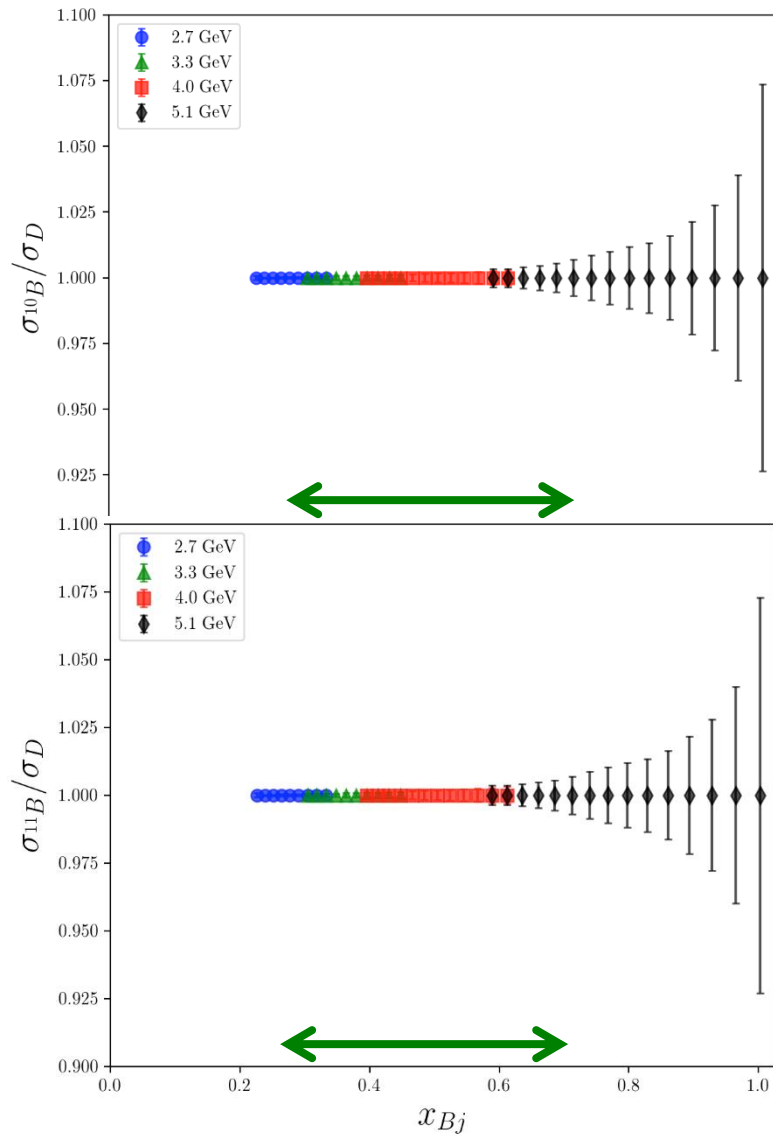
- Good coverage in  $x$  ( $0.15 < x < 0.9$ )
- Sizable amount of overlap in order to understand the acceptance
- Nominal  $x$  binning (0.025) provides good statistics over a wide  $x$  range
  - $< 1\%$  statistical error

*Stolen from E. Pooser*





# A dream of Spring ....2018 data



*New EMC (and SRC) data on*

- $^{12}\text{C}$
- $^{11}\text{B}$
- $^{10}\text{B}$
- $^9\text{Be}$

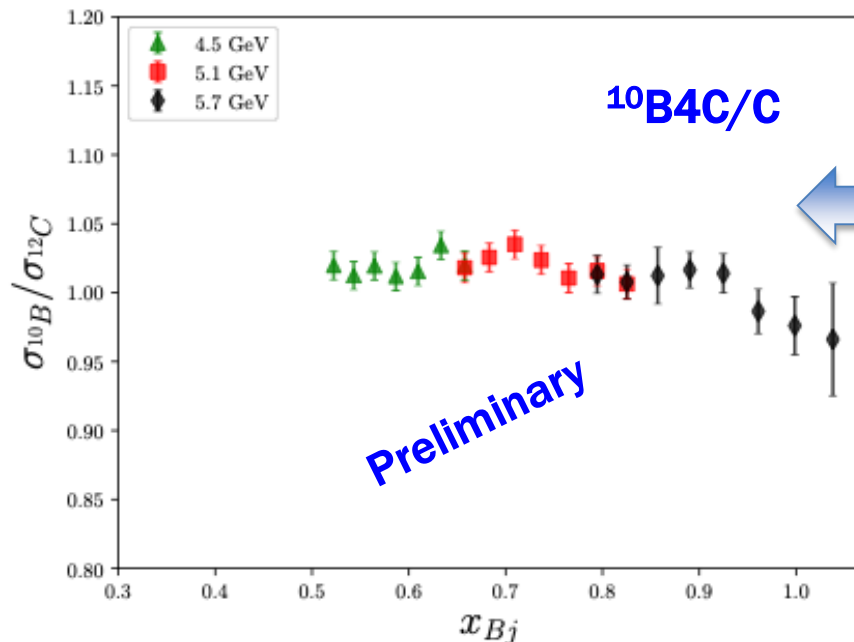
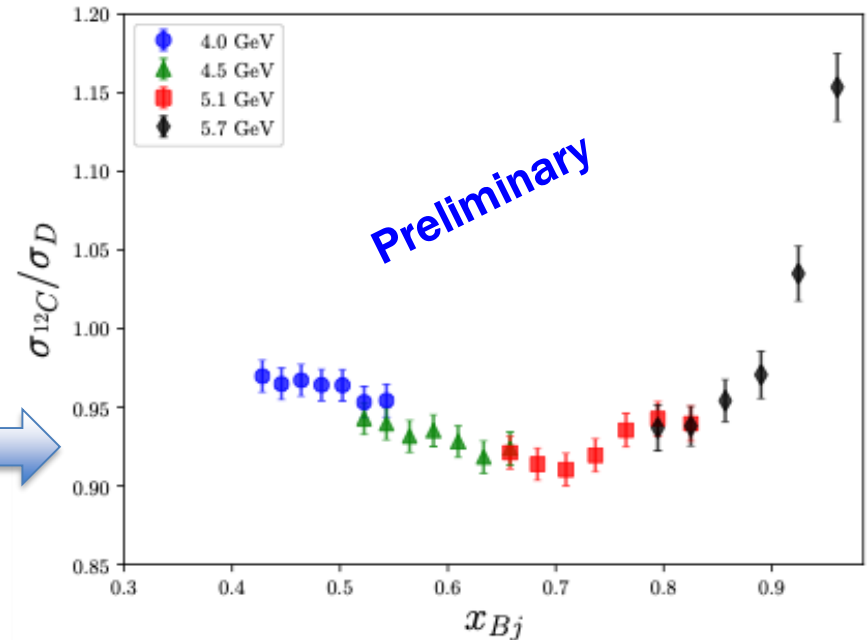
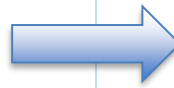
*Generously contributed by E. Pooser*

# Super-Preliminary Ratios: Spring 2018

## Preliminary ratios from HMS

- Larger  $x$  coverage and similar statistics in SHMS
- Uncertainties = statistics +  $\sim 1\%$  systematics

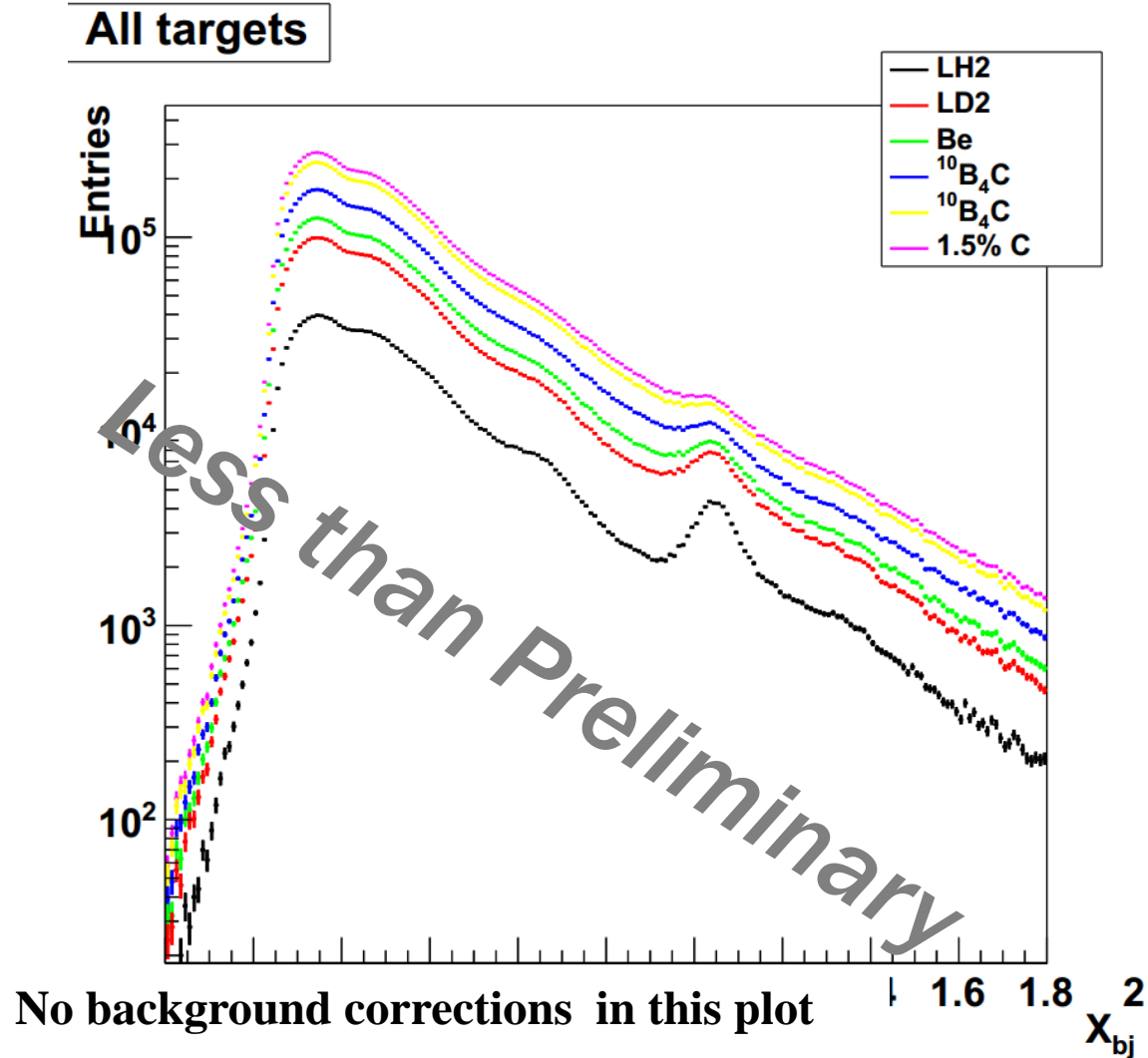
1. No bin-centering
2. First order acceptance correction (long vs. short target)
3. 0<sup>th</sup> order radiative correction



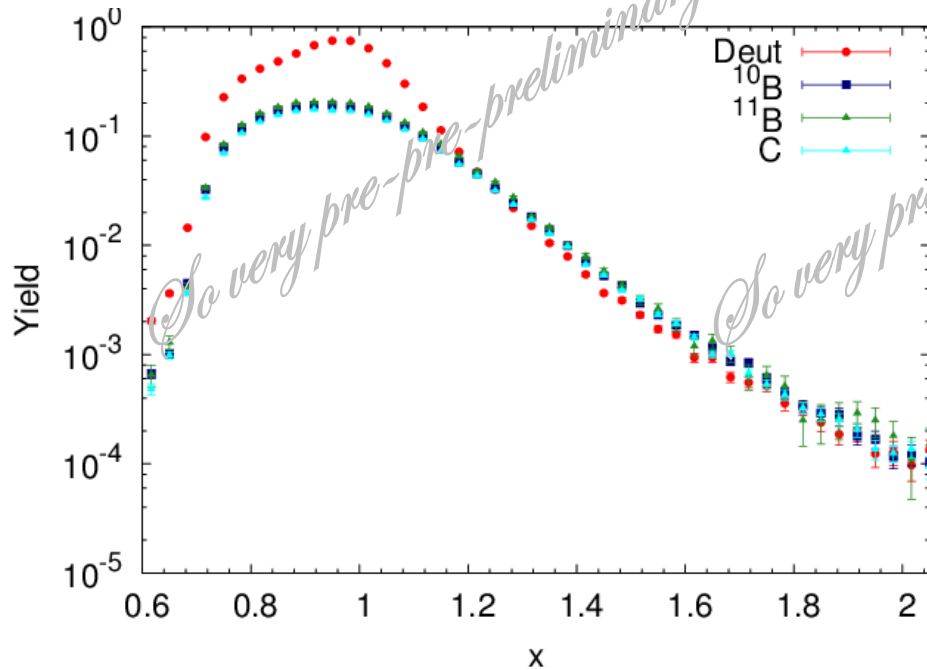
1. Ratio of raw yields (per nucleon)
2. No bin-centering
3. No acceptance correction (both point targets)
4. No radiative corrections

*Ratios by Eric Pooser*

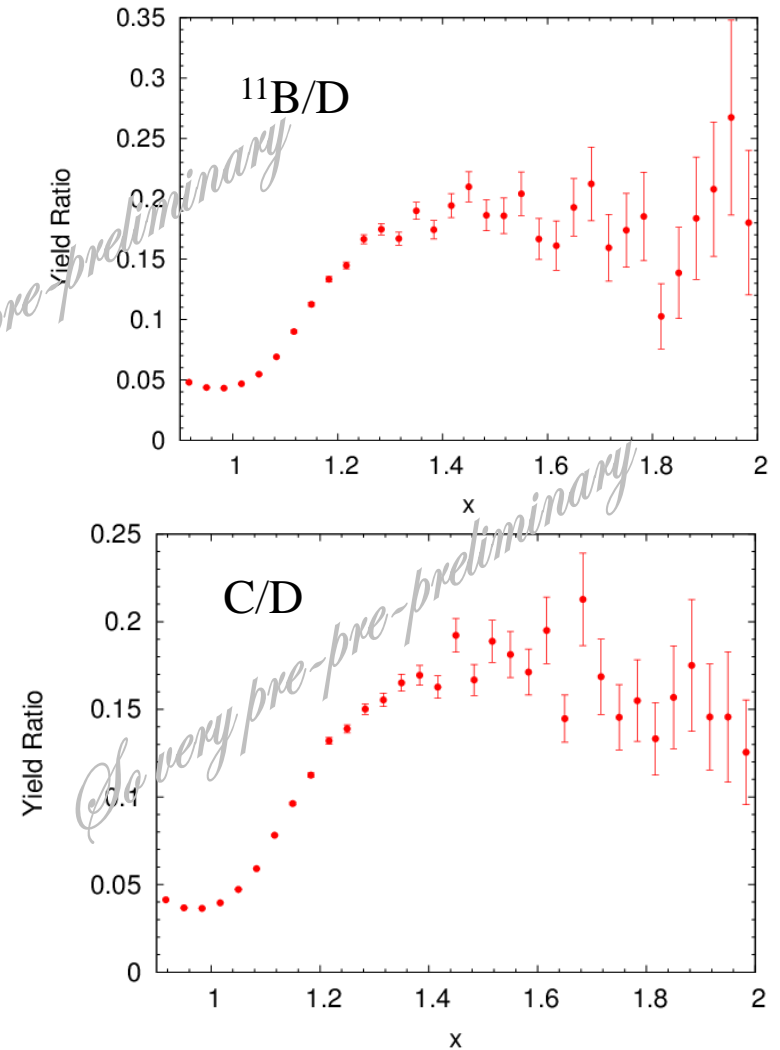
# And on the other side of $x=1$ (2018)



# Tried again in 2019

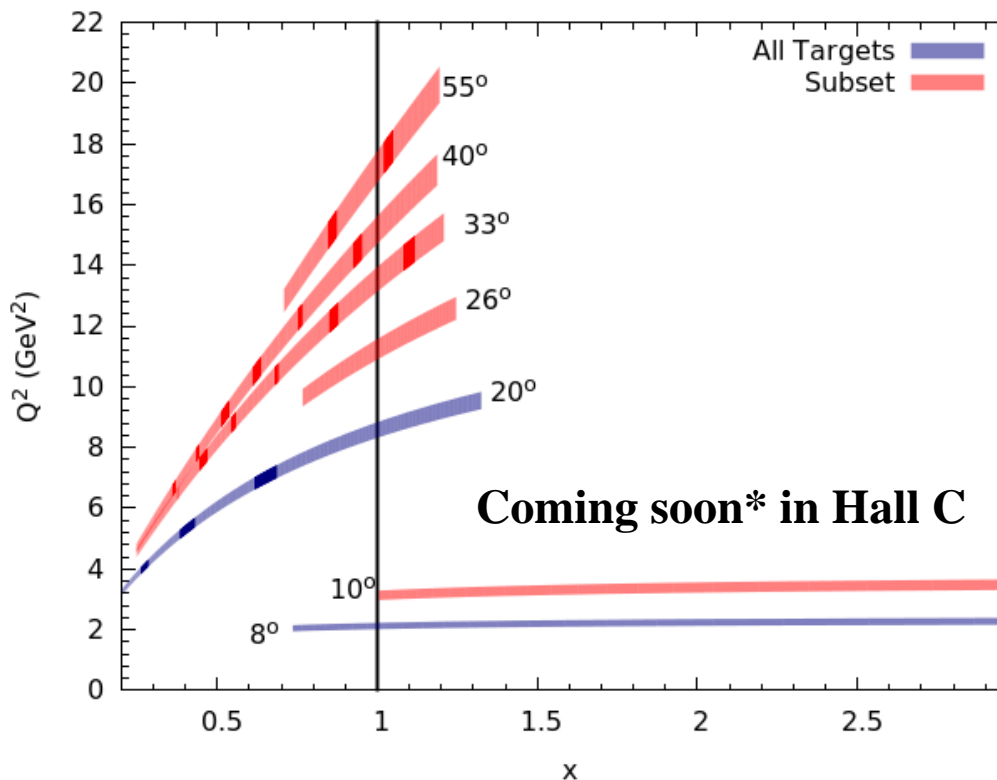


*Analysis in progress by Casey Morean*

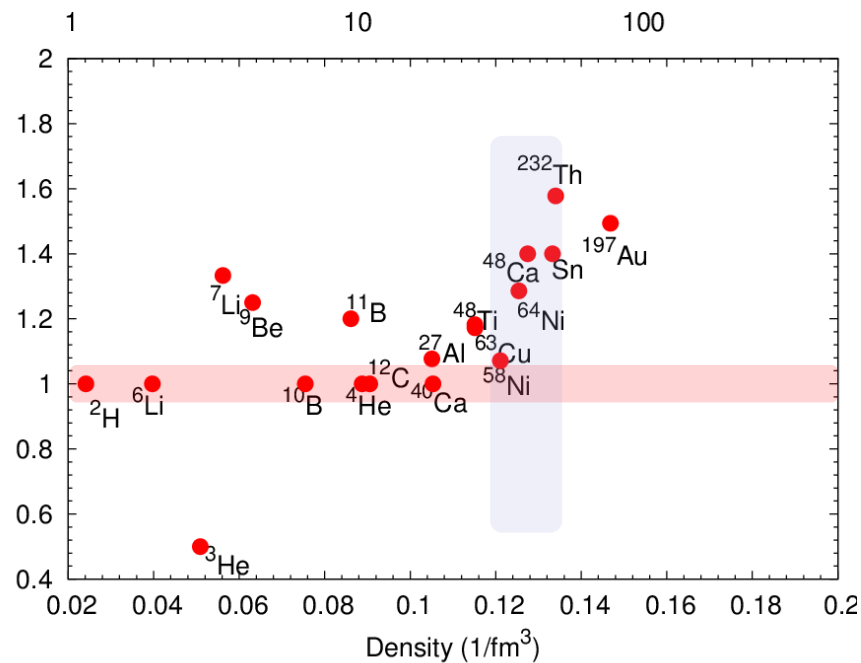
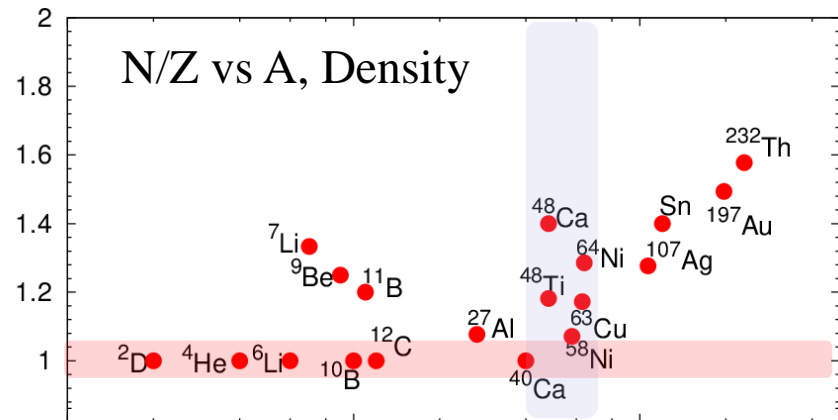


# What about all the other targets?

## Target Choice motivated by physics impact



- *Light nuclei – sensitivity to some nuclear structure (clusters, halos, etc...)*
- *Heavier nuclei – N/Z isospin lever*



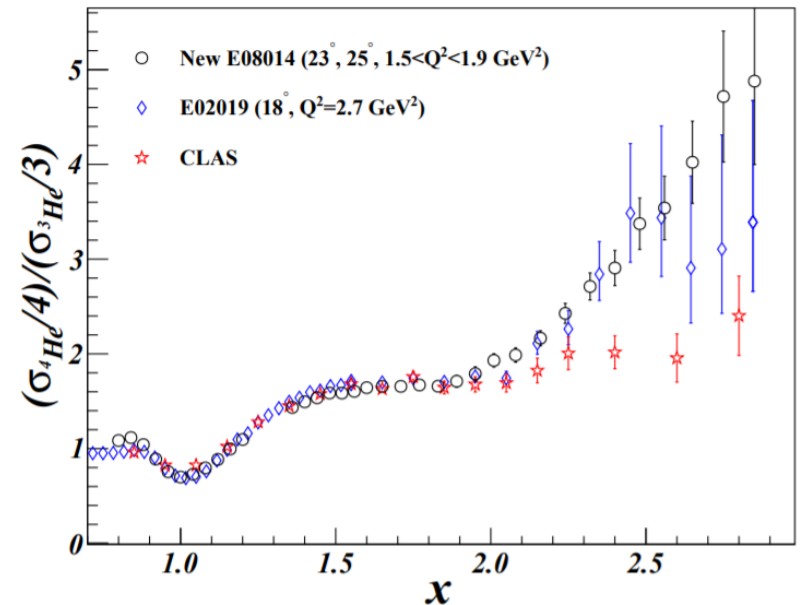
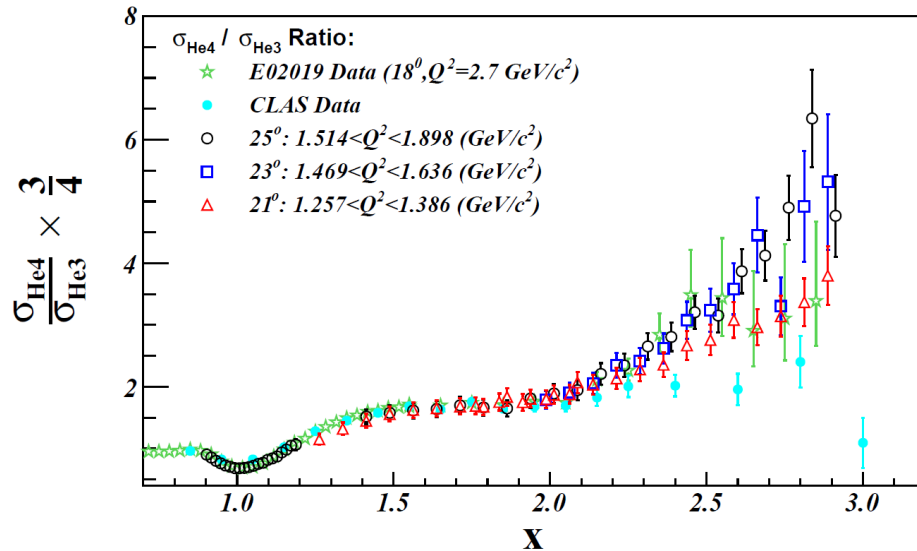
*What is soon?*

# 3N correlations - still looking

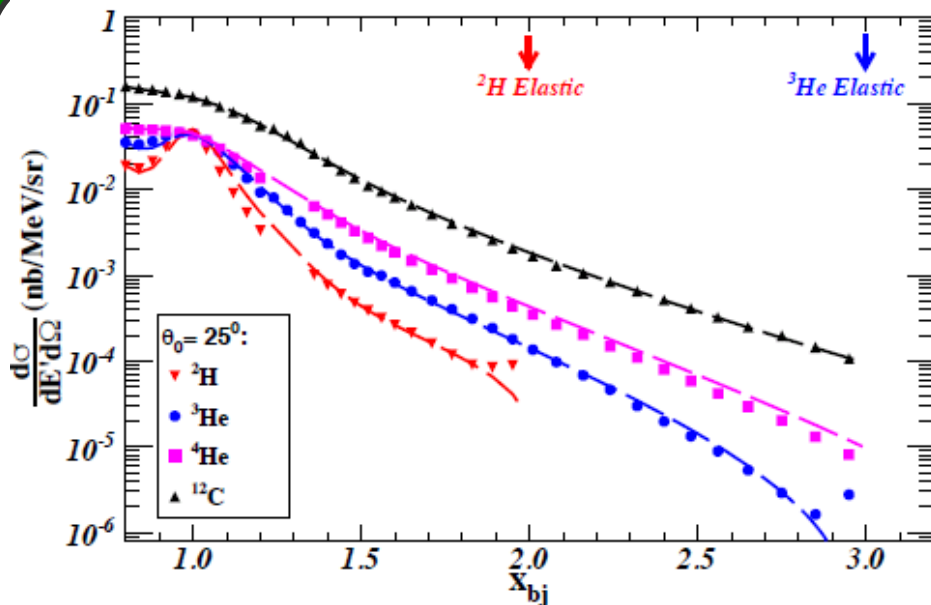
## Search for three-nucleon short-range correlations in light nuclei

Z. Ye,<sup>1,2,3</sup> P. Solvignon,<sup>4,5,\*</sup> D. Nguyen,<sup>2</sup> P. Aguilera,<sup>6</sup> Z. Ahmed,<sup>7</sup> H. Albataineh,<sup>8</sup> K. Allada,<sup>5</sup> B. Anderson,<sup>9</sup> D. Anez,<sup>10</sup> K. Aniol,<sup>11</sup> J. Annand,<sup>12</sup> J. Arrington,<sup>1</sup> T. Averett,<sup>13</sup> H. Baghdasaryan,<sup>2</sup> X. Bai,<sup>14</sup> A. Beck,<sup>15</sup> S. Beck,<sup>15</sup> V. Bellini,<sup>16</sup> F. Benmokhtar,<sup>17</sup> A. Camsonne,<sup>5</sup> C. Chen,<sup>18</sup> J.-P. Chen,<sup>5</sup> K. Chirapatpimol,<sup>2</sup> E. Cisbani,<sup>19</sup> M. M. Dalton,<sup>2,5</sup> A. Daniel,<sup>20</sup> D. Day,<sup>2</sup> W. Deconinck,<sup>21</sup> M. Defurne,<sup>22</sup> D. Flay,<sup>23</sup> N. Fomin,<sup>24</sup> M. Friend,<sup>25</sup> S. Frullani,<sup>19</sup> E. Fuchey,<sup>23</sup> F. Garibaldi,<sup>19</sup> D. Gaskell,<sup>5</sup> S. Gilad,<sup>21</sup> R. Gilman,<sup>26</sup> S. Glamazdin,<sup>27</sup> C. Gu,<sup>2</sup> P. Guèye,<sup>18</sup> C. Hanretty,<sup>2</sup> J.-O. Hansen,<sup>5</sup> M. Hashemi Shabestari,<sup>2</sup> O. Hen,<sup>28</sup> D. W. Higinbotham,<sup>5</sup> M. Huang,<sup>3</sup> S. Iqbal,<sup>11</sup> G. Jin,<sup>2</sup> N. Kalantarians,<sup>2</sup> H. Kang,<sup>29</sup> A. Kelleher,<sup>21</sup> I. Korover,<sup>28</sup> J. LeRose,<sup>5</sup> J. Leckey,<sup>30</sup> R. Lindgren,<sup>2</sup> E. Long,<sup>9</sup> J. Mammei,<sup>31</sup> D. J. Margaziotis,<sup>11</sup> P. Markowitz,<sup>32</sup> D. Meekins,<sup>5</sup> Z. Meziani,<sup>23</sup> R. Michaels,<sup>5</sup> M. Mihovilovic,<sup>33</sup> N. Muangma,<sup>21</sup> C. Munoz Camacho,<sup>34</sup> B. Norum,<sup>2</sup> Nuruzzaman,<sup>35</sup> K. Pan,<sup>21</sup> S. Phillips,<sup>4</sup> E. Piasetzky,<sup>28</sup> I. Pomerantz,<sup>28,36</sup> M. Posik,<sup>23</sup> V. Punjabi,<sup>37</sup> X. Qian,<sup>3</sup> Y. Qiang,<sup>5</sup> X. Qiu,<sup>38</sup> P. E. Reimer,<sup>1</sup> A. Rakhman,<sup>7</sup> S. Riordan,<sup>2,39</sup> G. Ron,<sup>40</sup> O. Rondon-Aramayo,<sup>2</sup> A. Saha,<sup>5,\*</sup> L. Selvy,<sup>9</sup> A. Shahinyan,<sup>41</sup> R. Shneur,<sup>28</sup> S. Sirca,<sup>42,33</sup> K. Slifer,<sup>4</sup> N. Sparveris,<sup>23</sup> R. Subedi,<sup>2</sup> V. Sulkosky,<sup>21</sup> D. Wang,<sup>2</sup> J. W. Watson,<sup>9</sup> L. B. Weinstein,<sup>8</sup> B. Wojtsekhowski,<sup>5</sup> S. A. Wood,<sup>5</sup> I. Yaron,<sup>28</sup> X. Zhan,<sup>1</sup> J. Zhang,<sup>5</sup> Y. W. Zhang,<sup>26</sup> B. Zhao,<sup>13</sup> X. Zheng,<sup>2</sup> P. Zhu,<sup>43</sup> and R. Zielinski<sup>4</sup>

(The Jefferson Lab Hall A Collaboration)

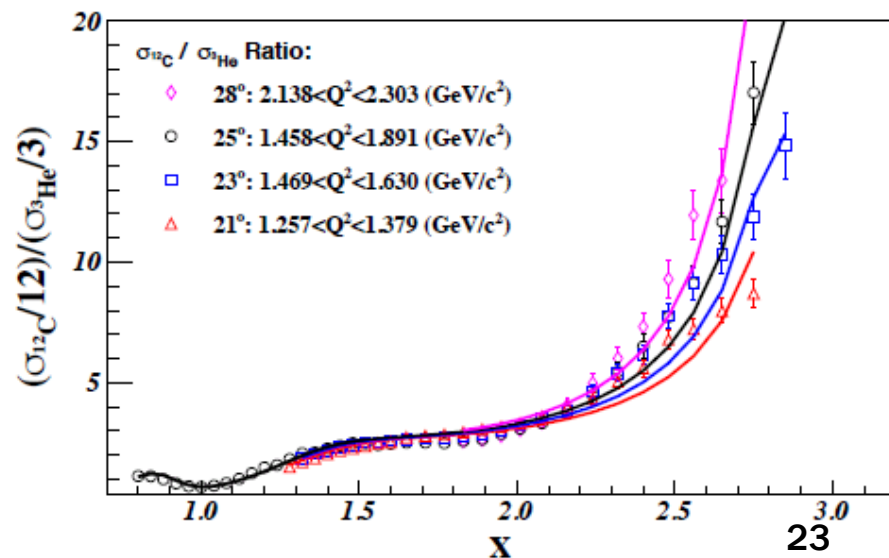
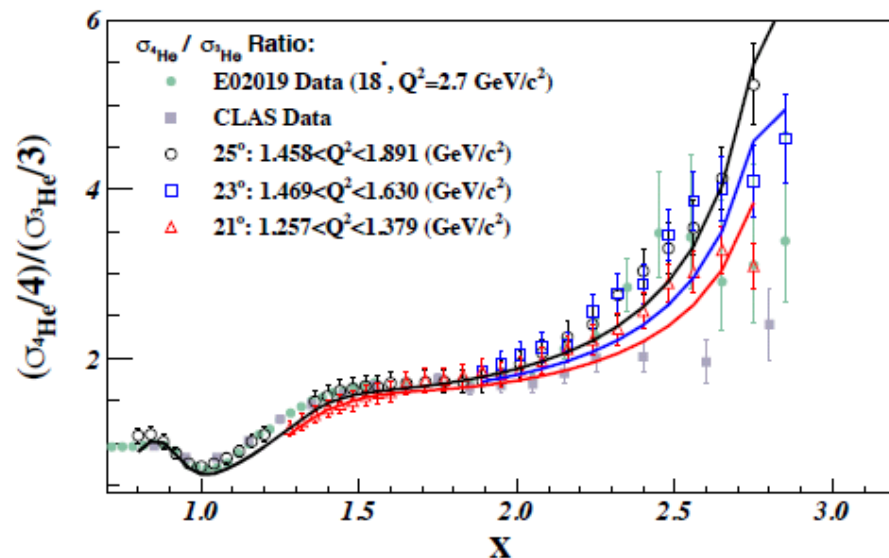


# Can we see a second plateau?



**Deuteron: smeared SRC similar to  ${}^2\text{H}$**   
(A/D is ~flat) until  $x > 1.8$

**${}^3\text{He}$ : cross section of stationary 3N-SRC begins to fall off closer to  $x = 2.6$ . Sets in EARLIER at high  $Q^2$**



**I WAS TOLD THERE WOULD BE**



**3N SRC PLATEAUS**





# Summary

- New inclusive data on EMC effect and SRC ratios at 12 GeV on new nuclei
- Looking forward to data on:
  - Additional light nuclei (extra neutrons!)
  - Scans in n/p at fixed A
  - Pushing to  $x=3$  looking for 3N SRCs
- Results on new nuclei soon\*!