# Physics GRE 1777 Solutions Mohamed Abdelhafez abdelhafez@uchicago.edu

# September 5, 2017

Contents	19 CM: Energy
1 CM D	20 EM: Sources of B
1 CM: Forces	21 EM: Magnetic Force
2 CM: Momentum	22 AP: Bohr Model
3 CM: SHM	23 AP: Bohr Model
4 EM: Circuits	24 AP: Electronic Configuration
5 EM: Maxwell	25 QM: Operators
6 EM: EM Waves	26 QM: Commutation Relations
7 TD: Types of Processes	27 ST: Nuclear Physics
8 TD: Types of Processes	28 TD: Cycles
9 TD: Distribution of Speeds	29 TD: Types of Processes
10 ST: Photoelectric Effect	30 WO: Doppler Effect
11 ST: Characteristic X-Rays	31 WO: Standing Waves
12 QM: Angular Momentum	32 WO: Telescopes
13 QM: Particle in a Box	33 ST: Lasers
14 LM: Dimensional Analysis	34 SR: Length Contraction
15 WO: Refraction	35 SR: Time Dilation
16 WO: Refraction	36 SR: Energy-Momentum
17 CM: Orbits	
	37 EM: Gauss's Law
18 CM: Orbits	38 EM: Potential

39 CM: Forces	66 CM: Springs
40 EM: Circuits	67 CM: Hamiltonians
41 LM: Error Analysis	68 EM: Gauss's Law
42 LM: Error Analysis	69 EM: Induction
43 QM: Hydrogen Atom	70 QM: Incidence on V
44 QM: States	71 QM: Operators
45 QM: Incidence on V	72 ST: Nuclear Physics
46 CM: Springs	73 ST: Particle Physics
47 CM: Fluid Dynamics	74 ST: Particle Physics
48 CM: Rigid Bodies	75 ST: Bragg's Law
49 CM: Projectile	76 EM: Magnetic Force
50 WO: Lenses	77 EM: Boundary Conditions
51 WO: Resolution	78 EM: Radiation
52 WO: Interferometer	79 CM: Forces
53 SR: Doppler Shift	80 CM: Gravitation
54 LM: Electronics	81 WO: Thin Films
55 ST: Fourier Series	82 WO: Refraction
56 TD: Distribution of Speeds	83 ST: Radiation Power
57 TD: Distribution of Speeds	84 SR: Doppler Shift
58 TD: Entropy	85 SR: Energy-Momentum
59 ST: Frank Hertz	86 ST: Hall Effect
60 ST: Compton Effect	87 CM: Energy
61 QM: Pauli Exclusion Principle	88 CM: Relative Velocity
62 QM: Harmonic Oscillator	89 CM: Terminal Speed
63 QM: Angular Momentum	90 CM: SHM
64 CM: Energy	91 TD: Carnot
65 CM: Rigid Bodies	92 QM: Particle in a Box

93 SR: 4-Vectors 97 CM: Lagrangians

94 EM: Circuits 98 ST: Bragg's Law

95 EM: Potential 99 QM: Selection Rules

96 EM: Sources of B 100ST: Mössbauer Spectroscopy

## Abbreviations

CM: Classical Mechanics

EM: Electricity and Magnetism

WO: Waves and Optics

AP: Atomic Physics

SR: Special Relativity

QM: Quantum Mechanics

TD: Thermodynamics

LM: Lab Methods

ST: Special Topics

### 1 CM: Forces

Answer: (E)

 $F_A = m_A a_A, \ F_B = m_B a_B = 2 \ m_A 2 \ a_A = 4 F_A.$ 

#### 2 CM: Momentum

Answer: (E)

$$\vec{P}_i = \vec{P}_f \implies 0.5 * 2\hat{i} + 1 * 1\hat{j} = 1.5\vec{v}_f$$
. So,  $\vec{v}_f = \frac{1}{1.5}(\hat{i} + \hat{j}) \implies v_f = \frac{2\sqrt{2}}{3}$  m/s. So,  $KE_f = \frac{1}{2}(1.5)(\frac{2\sqrt{2}}{3})^2 = \frac{2}{3}$  J, and  $KE_i = \frac{1}{2}(0.5)(2)^2 + \frac{1}{2}(1)(1)^2 = \frac{3}{2}$  J. So  $KE_i - KE_F = \frac{3}{2} - \frac{2}{3} = \frac{9-4}{6} = \frac{5}{6}$  J.

### 3 CM: SHM

Answer: (E)

$$\omega = \sqrt{\frac{g}{L}}$$
, so  $T = \frac{2\pi}{\omega} = 2\pi \frac{L}{g}$ . Therefore,  $\frac{T_A}{T_B} = \sqrt{\frac{L_A}{L_B}} \implies \frac{L_A}{L_B} = \frac{1}{4} \implies L_B = 4L_A$ .

### 4 EM: Circuits

**Answer:** (B) We can combine the two rightmost resistors in parallel into  $R_{eq} = \frac{1*2}{1+2} = \frac{2}{3}\Omega$ . So the total current in the circuit is  $I_{tot} = \frac{20}{1+\frac{2}{3}} = 12A$  from V = IR. Then we wanna divide this current between two branches for the 1 and 2  $\Omega$ . We can do that by knowing they share the same potential:  $I_1 * 1 = I_2 * 2$  and  $I_1 + I_2 = 12$ , leading to  $2I_2 + I_2 = 12 \implies I_2 = 4A$ .

Alternatively, we could write Kirchoff's law for the two loops:

$$-20 + I_1(1) + (I_1 - I_2)(1) = 0$$
 and  $(I_2 - I_1)(1) + 2I_2 = 0$  and solve for  $I_2$ :  $3I_2 = I_1 \implies -20 + 6I_2 - I_2 = 0 \implies I_2 = 4A$ .

#### 5 EM: Maxwell

Answer: (E)

Remember one of Maxwell's equations  $\vec{\nabla} \times \vec{B} = \mu_0(\vec{J} + \epsilon_0 \frac{\partial \vec{E}}{\partial t})$ , where the second term is the displacement current density. So we see that if we integrate over the area of S, the displacement current will be proportional to the rate of change of the electric flux through S.

#### 6 EM: EM Waves

Answer: (B)

In EM waves,  $\vec{E}$  is perpendicular to  $\vec{B}$  and they are both perpendicular to the direction of propagation. Here the direction of propagation is  $+\hat{e_z}$  (from  $\sin(kz - \omega t)$ ), so the only choice that is perpendicular to both  $\hat{e_z}$  and to  $(\hat{e_x} + \hat{e_y})$  is  $(-\hat{e_x} + \hat{e_y})$ . (Exact direction can be calculated from  $\vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$ ).

5

## 7 TD: Types of Processes

Answer: (C)

A reversible process must have  $\Delta S = 0$ .

### 8 TD: Types of Processes

Answer: (B)

In general,  $\Delta U = Q - W$ . So if  $\Delta U = Q$ , then W = 0. But  $W = \int P dV$ , so  $\Delta V = 0$ .

### 9 TD: Distribution of Speeds

Answer: (D)

 $v_{rms} = \sqrt{\frac{3kT}{m}} = \sqrt{\frac{3RT}{M}}$  is one of the main equations you should know for the GRE.

### 10 ST: Photoelectric Effect

**Answer:** (D) The equation for the photoelectric effect is  $h\nu = \phi + KE$  if  $\nu > \nu_{cuttoff}$ , where  $\phi$  is the metal work function. And then the stopping potential should cancel the KE, so  $e|V_{stop}| = h\nu - \phi$ .

### 11 ST: Characteristic X-Rays

Answer: (C)

Characteristic X-Rays are emitted when outer-shell electrons fill a vacancy in the inner shell of the atom.

# 12 QM: Angular Momentum

Answer: (E)

For any l, there are (2l+1) possible values for m from -l to +l. So we have 5 possible m values here: -2, -1, 0, 1, 2.

# 13 QM: Particle in a Box

Answer: (D)

The energies for a particle in a box are  $E_n = \frac{n^2 \hbar^2 \pi^2}{2ma^2}$ , so  $E_1 = \frac{\hbar^2 \pi^2}{2ma^2}$ .

# 14 LM: Dimensional Analysis

Answer: (A)

The only choice with the correct units is (A). This is a repeated question from one of the previous GRE exams.

### 15 WO: Refraction

Answer: (C)

The index of refraction is related to the dielectric constant by  $n = \sqrt{k}$ . So, n = 2, therefore  $v = \frac{c}{n} = 1.5 \times 10^8$  m/s.

#### 16 WO: Refraction

Answer: (D)

Fermat's principle tells us how light chooses a path to move between two points, so it is how we prove how light moves when hits a boundary, like in reflection and refraction.

#### 17 CM: Orbits

Answer: (D)

From Kepler's third law:  $T^2 \propto r^3$ , so,  $(\frac{T_1}{T_2})^2 = (\frac{r_1}{r_2})^3 = (\frac{1}{4})^3 = \frac{1}{64}$ . Therefore,  $\frac{T_1}{T_2} = \frac{1}{8}$ .

#### 18 CM: Orbits

Answer: (C)

For circular orbits,  $\frac{mv^2}{R} = \frac{GmM}{R^2} \implies v = \sqrt{\frac{GM}{R}}$ . But the angular momentum L = mvR, so L =  $m\sqrt{GMR}$  and  $\frac{L_A}{L_B} = \sqrt{\frac{R_A}{R_B}} = \sqrt{2}$ .

### 19 CM: Energy

Answer: (C)

 $E_i + W = E_f \implies \frac{1}{2}mv_i^2 + Fd = \frac{1}{2}mv_f^2$ . So,  $\frac{1}{2}(10)((2)^2 - (1)^2) = 5F \implies 15 = 5F \implies F = 3N$ 

### 20 EM: Sources of B

Answer: (D)

One of the equations to know for the GRE is the magnetic field of a current loop at the center (can be derived from Biot-Savart Law)  $B_{center} = \frac{\mu_0 I}{2B}$ .

### 21 EM: Magnetic Force

Answer: (E)

For cyclotrons,  $\frac{mv^2}{R} = qvB \implies m\omega = qB$ , so  $\omega = \frac{qB}{m}$ . For a proton, q = e,  $m = m_p$ . For a deutron, q = e,  $m = 2m_p$ . For an alpha particle, q = 2e,  $m = 4m_p$ . Therefore,  $\omega_p : \omega_d : \omega_\alpha = 1 : \frac{1}{2} : \frac{1}{2}$ . Therefore,  $\omega_p > \omega_d = \omega_\alpha$ . (Here, we approximate the mass of the neutron to be equal to the mass of the proton).

7

#### 22 AP: Bohr Model

### Answer: (A)

In the Bohr model,  $E_n \propto Z^2$ . Therefore,  $\frac{E_{Li}}{E_H} = \frac{3^2}{1^2} = 9$ . But  $E \propto f \propto \frac{1}{\lambda} \implies \frac{\lambda_{Li}}{\lambda_H} = \frac{1}{9}$ .

#### 23 AP: Bohr Model

#### Answer: (B)

Bohr model has E  $\propto \mu$  (the reduced mass). For the positronium, the reduced mass is  $\mu =$  $\frac{m_e m_e}{m_e + m_e} = \frac{m_e}{2}$  which is just half the mass for the Hydrogen atom. Therefore,  $\frac{E_{pos}}{E_H} = \frac{1}{2}$ .

#### **AP: Electronic Configuration** 24

#### Answer: (C)

The Nitrogen atom has the electronic configuration of  $1s^2$   $2s^2$   $2p^3$ . The three electrons in  $2p^3$  all have parallel spins since p has a capacity of 6, and when filling it, you fill it with three parallel spins first then you couple the extra electrons in an anti-parallel sense (known as Hund's Rule). Therefore, the total spin number for the Nitrogen atom is  $3 \times \frac{1}{2} = \frac{3}{2}$ .

#### 25 QM: Operators

#### Answer: (B)

For hermitian operators, the eigenvalues are real. So only choices (A) and (B) survive. Also, the eigenvalues have to satisfy  $a^4 = 1$ , so answer must be (B).

#### 26 QM: Commutation Relations

**Answer: (E)**  $[T, P] = [\frac{P^2}{m}, P] = 0$  since  $[P^2, P] = P^2P - PP^2 = P^3 - P^3 = 0$ . So, T and P are compatible observables. All the other choices will result in commutators that relate position to momentum, which will not be zero because X, P are incompatible observables,  $[X, P] = i\hbar$ .

#### 27 ST: Nuclear Physics

#### Answer: (A)

Well, we know that nuclei decay via either  $\alpha^{+2}$ ,  $\beta^{+1}$  or  $\gamma$ . So,  $\gamma$  rays are a very likely product from a nucleus.

#### 28 TD: Cycles

#### Answer: (C)

In a cycle,  $\Delta U = 0$  so Q = W = area enclosed by the cycle. Here, we have a clockwise cycle which is the condition for positive work and heat. So,  $Q = W = \frac{1}{2}(6000 - 2000)(0.03 - 0.01) = 40J$ .

### 29 TD: Types of Processes

#### Answer: (D)

1 has constant pressure so it's isobaric. And we (should) know that the work done by an isothermal process is always bigger than the work done by an adiabatic process starting from the same state ( $W_{isothermal} > W_{adiabatic}$ ). Therefore 2 must be the isothermal process, 3 the adiabatic.

### 30 WO: Doppler Effect

#### Answer: (E)

Here we have a double Doppler shift because first the wall receives a shifted frequency from the source (the moving car), then it reflects it and now the moving car becomes the observer and it's moving. So in total  $f' = \frac{v+v_0}{v-v_0} f_0 = \frac{350+3.5}{350-3.5} 600 = \frac{1+0.01}{1-0.01} 600$ .

A useful approximation you might use and it will always make your life easier in Doppler

A useful approximation you might use and it will always make your life easier in Doppler problems (especially relativistic ones) is that  $\frac{1}{1-x} \approx 1+x$  if  $x \ll 1$ . Hence,  $f' \approx (1+0.01)^2$  600 =  $(1+0.02+0.0001) \times 600 \approx 1.02*600 = 612$  Hz.

### 31 WO: Standing Waves

#### Answer: (E)

For open-open, we have  $f_m = \frac{mv}{2L}$ . Using one of the given modes,  $50 = \frac{v}{2L} \implies \frac{v}{L} = 100$ . So if we close one end, the modes become  $f_m = \frac{mv}{4L}$  where m is odd. So,  $f_m = \frac{m100}{4} = 25m = 25$ , 75, 125, 175, ... So none survive!

# 32 WO: Telescopes

#### Answer: (A)

For a telescope:  $f_o + f_e = L$  and  $|M| = \frac{f_0}{f_e}$ . So  $20 + f_o = 100 \implies f_0 = 80$  cm. Therefore,  $M = \frac{80}{20} = 4$ .

#### 33 ST: Lasers

#### Answer: (A)

Dye lasers are very tunable to cover the visible wavelength spectrum.

# 34 SR: Length Contraction

### Answer: (B)

 $L = \frac{L_0}{\gamma} \implies \gamma = \frac{1}{0.8} \implies \beta = 0.6$ . (Remember the fact that  $\beta$  and  $\frac{1}{\gamma}$  are Pythagorean pairs with a hypotenuse of 1).

#### 35 SR: Time Dilation

#### Answer: (D)

 $\beta = 0.6 \implies \frac{1}{\gamma} = 0.8$ . Therefore,  $\Delta t = \gamma \Delta t_0 = \frac{2 \times 10^{-6}}{0.8}$ . So,  $d = vt = 0.6 \times 3 \times 10^8 \times \frac{2 \times 10^{-6}}{0.8} = \frac{9}{2} \ 100 = 450m$ .

### 36 SR: Energy-Momentum

#### Answer: (D)

 $m^2 = E^2 - p^2$  (setting c = 1). So,  $m^2 = (5)^2 - (5 - 0.1)^2 = (5)^2 - [(5)^2 + (0.1)^2 - 2 \times 5 \times 0.1] = 1 - 0.01 \approx 1$ . So, m = 1 GeV/ $c^2$ .

#### 37 EM: Gauss's Law

### Answer: (D)

The charge makes a total flux around it of  $\frac{Q}{\epsilon_0}$  by Gauss's law. But the charge is above the yz plane and half the field lines produced by the charge will go down and hit the yz plane while the other half will go up and will not be felt by the yz plane. So, the total flux through the plane is half the total flux of the charge  $=\frac{Q}{2\epsilon_0}$ .

### 38 EM: Potential

#### Answer: (E)

Our reference for V = 0 is the usual reference and it allows us to think of the whole setup from outside as a sphere potential, so  $V(r>b)=\frac{kQ_{tot}}{r}=\frac{k(Q+q)}{r}$ . And potential should be continuous, so V at the boundary of r = b is equal to  $V(b)=\frac{k(Q+q)}{b}$ . But the field inside the conductor is zero, and hence the potential should be constant between r=a and r=b, so in this whole region we just have  $V(a< r< b)=\frac{k(Q+q)}{b}$ .

### 39 CM: Forces

#### Answer: (C)

- 1.  $F_1 = F_d + F_D$ .
- 2.  $F_2 = F_d F_D$ .
- 3.  $F_3 = \sqrt{F_d^2 + F_D^2}$ .

(here  $F_D$  and  $F_d$  are just the positive magnitudes). Clearly,  $F_2$  is the smallest. We know that  $F_d + F_D > \sqrt{F_d^2 + F_D^2}$  since if we square the LHS, we get  $(F_d + F_D)^2 = F_d^2 + F_D^2 + 2F_DF_d > F_d^2 + F_D^2$ . Therefore, we see that  $F_1 > F_3 > F_2$ .

#### 40 EM: Circuits

#### Answer: (A)

The only element that dissipates energy is R. So, we should expect a result  $P \propto I^2 R$ . The exact equation for power in AC is  $P = V_{rms}I_{rms}cos\phi$ . But  $V_{rms} = I_{rms}Z$ . So,  $P = I_{rms}^2 Z cos\phi$ . But  $Z cos\phi$  is just R. (Remember  $Z = \sqrt{R^2 + (X_L - X_C)^2}$  and  $\tan \phi = \frac{X_L - X_C}{R}$ , so if you construct that right angle triangle of R,  $(X_L - X_C)$  and Z, you see that  $Z cos\phi = R$ .). Hence  $P = I_{rms}^2 R$  which we could have guessed anyway.

### 41 LM: Error Analysis

#### Answer: (E)

The average number of photons N = 0.1 (100) = 10. A decaying process follows Poisson distribution where  $\sigma = \sqrt{N} = \sqrt{10} \approx 3$ . So N = 10 ± 3.

### 42 LM: Error Analysis

#### Answer: (D)

Given  $\frac{\sigma_v}{v} = 0.1$ , we want to calculate  $\frac{\sigma_f}{f}$  where  $f = \frac{1}{2}mv^2$ . Applying the propagation of error formula,  $\sigma_f = \frac{\partial f}{\partial v}\sigma_v = mv\sigma_v = \frac{mv^2 \sigma_v}{v} = 0.1mv^2$ . Therefore,  $\frac{\sigma_f}{f} = \frac{\sigma_f}{0.5mv^2} = \frac{0.1}{0.5} = 0.2$ .

### 43 QM: Hydrogen Atom

#### Answer: (C)

2s is spherically symmetric, so there should be no dependence on  $\theta$  or  $\phi$ . Only choices (B), (C) remain. But (B) is the ground state of the hydrogen atom (1s). So the answer must be (C).

### 44 QM: States

### Answer: (A)

The normalization condition is  $\sum \text{Prob} = 1 \implies A^2 + A^2 + A^2 = 1 \implies A = \frac{1}{\sqrt{3}}$ .

### 45 QM: Incidence on V

### Answer: (A)

Remember that  $\frac{\hbar^2 k^2}{2m} = E - V \implies k = \frac{1}{\hbar} \sqrt{2m(E - V)}$ . Since regions 1 and 3 have the same E and V, they will definitely have the same k. There is also no reason for the solution to have the special forms of sines or cosines, so generally it would be just  $Ae^{ikx}$ . The incident wave oscillates with k in region 1, oscillates with a different k' in region 2, then back to oscillating with k in region 3.

# 46 CM: Springs

#### Answer: (A)

F =  $k\Delta x$  and  $\Delta U = \frac{1}{2}k\Delta x^2 = \frac{1}{2}F\Delta x$ . So  $\frac{\Delta x_1}{\Delta x_2} = \frac{k_1}{k_2} \implies \Delta x_1 < \Delta x_2$ . And  $\frac{\Delta U_1}{\Delta U_2} = \frac{\Delta x_1}{\Delta x_2} \implies \Delta U_1 < \Delta U_2$ .

# 47 CM: Fluid Dynamics

#### Answer: (A)

Before:  $mg = \rho_W g(0.5V_B)$ . After:  $mg = \rho_W g(V_{stone} + V_B')$ . Since mg is the same, we see that  $0.5 \ V_B = V_{stone} + V_B'$ . Therefore,  $V_B' < 0.5V_B$ .

### 48 CM: Rigid Bodies

#### Answer: (B)

$$I_D = \frac{1}{2}MR^2$$
. So,  $KE_{rot} = \frac{1}{2}I\omega^2 = \frac{1}{2}\frac{1}{2}MR^2\frac{v^2}{R^2} = \frac{1}{4}mv^2 = \frac{1}{2}KE_{trans}$ . Therefore,  $\frac{KE_{rot}}{KE_{tot}} = \frac{KE_{rot}}{KE_{rot}+KE_{trans}} = \frac{\frac{1}{2}}{\frac{1}{2}+1} = \frac{1}{3}$ .

# 49 CM: Projectile

#### Answer: (B)

Therefore, 
$$\frac{h_1}{h_2} = (\frac{\sin \theta_1}{\sin \theta_2})^2$$
. So,  $\frac{\sin \theta_1}{\sin \theta_2} = \sqrt{2}$ .

#### 50 WO: Lenses

#### Answer: (C)

For the second lens,  $\frac{1}{f} = \frac{1}{s} + \frac{1}{s'} \implies \frac{1}{20} = \frac{1}{10} + \frac{1}{s'} \implies s' = -20$  cm. So, it is a virtual image (since s' is negative) and it is inverted relative to the original object, but upright relative to the first image.

### 51 WO: Resolution

#### Answer: (B)

Resolving power is  $\frac{\lambda}{\Delta\lambda} = \frac{500}{502 - 500} = 250$ .

### 52 WO: Interferometer

#### Answer: (B)

As the gas is evacuated, the path difference for light is 2d (n-1) (full gas - empty gas). So, 2d (n-1) = m  $\lambda \implies 2(0.1)(n-1) = 100 \times 632.2 \times 10^{-9}$ . So n-1 = 316.1 ×10<sup>-6</sup>  $\implies n = 1.000316$ . (If they mean by the optical path length nd, then the above formula is missing a division by n, but as n is very close to 1, it won't really matter).

# 53 SR: Doppler Shift

#### Answer: (D)

From Hubble's law,  $v = H_0$  r,  $v = 75 \times 10^3 \times 100 = 75 \times 10^5$  m/s. So,  $\beta = \frac{v}{c} = 0.025$ . Then the Doppler shift follows  $\frac{\lambda'}{\lambda} = \sqrt{\frac{1+\beta}{1-\beta}} \implies (\frac{\lambda'}{\lambda})^2 = \frac{1+0.025}{1-0.025}$ . Using the usual approximation for Doppler shifts,  $(\frac{1}{1-x} \approx 1 + x \text{ if } x \ll 1)$ ,  $\lambda' \approx \sqrt{(1+0.025)^2}\lambda = 1.025\mu\text{m}$ . So 25 nm longer!

#### 54 LM: Electronics

#### Answer: (B)

A diode blocks currents in one direction and allows them in the other, hence the sine wave will just exist in its positive half-cycle.

### 55 ST: Fourier Series

#### Answer: (B)

f(x) is even (f(x) = f(-x)), so it can't have any sine terms in the expansion, since sine is odd. So,  $b_n = 0$  for all n.

### 56 TD: Distribution of Speeds

#### Answer: (B)

The total area should be equal to the number of molecules N. So  $\frac{1}{2}av_0 + av_0 + \frac{1}{2}av_0 = N \implies 2av_0 = N$  and  $a = \frac{N}{2v_0}$ .

### 57 TD: Distribution of Speeds

#### Answer: (D)

The distribution of speeds at temperature T has a non-zero peak and starts from zero at v=0 and then dies eventually again. Therefore, there is always zero probability of v=0 (III is correct) and the peak is not at v=0 (II is wrong). Also, the velocities at equilibrium should average to zero once you include the directions of the vectors (I is correct). Hence I and III are correct.

### 58 TD: Entropy

#### Answer: (A)

$$V_i = V_f \implies W = 0 \implies \Delta U = Q = nc_V \Delta T$$
. So,  $\Delta S = \int \frac{dQ}{T} = \int \frac{nc_V dT}{T} = nc_V \ln(\frac{T_f}{T_i})$ . And  $c_V = \frac{3}{2}R$ , hence  $\Delta S = \frac{3}{2}nR\ln(\frac{T_f}{T_i})$ .

#### 59 ST: Frank Hertz

### Answer: (C)

The Frank Hertz experiment gives peaks at multiples of one transition frequency. So 4.9 Hz,  $2 \times 4.9$  Hz,  $3 \times 4.9$  Hz,  $\cdots$  Note that no atom has evenly spaced levels anyway, so these evenly spaced energies are just an indicator of one transition.

# 60 ST: Compton Effect

#### Answer: (A)

The equation for the Compton effect is exactly choice (A),  $\lambda' = \lambda + \frac{h}{mc}(1 - \cos \theta)$ .

# 61 QM: Pauli Exclusion Principle

#### Answer: (B)

Pauli exclusion principle states that two fermions must have a totally antisymmetric combined wavefunction. So the para case has antisymmetric spin, but symmetric spatial wavefunction. While the ortho has symmetric spin and antisymmetric spatial wavefunction. And two identical particles are on average closer (higher in energy) if their spatial wavefunction is symmetric.

# 62 QM: Harmonic Oscillator

### Answer: (D)

For a 3D harmonic oscillator,  $E = (n_x + n_y + n_z + \frac{3}{2})\hbar\omega$ . So we need  $n_x + n_y + n_z = 2$  which can happen in six different ways: (2,0,0), (0,2,0), (0,0,2), (1,1,0), (1,0,1) and (0,1,1).

### 63 QM: Angular Momentum

#### Answer: (A)

 $[J^2, J_z] = 0$  since we know they are compatible and share simultaneous eigenstates | jm > such that  $J^2 | \text{ jm} > = \text{ j} (\text{j+1}) | \text{ jm} > \text{ and } J_z | \text{ jm} > = \text{ m} \hbar | \text{ jm} >$ .

### 64 CM: Energy

#### Answer: (E)

In free fall,  $v_f^2 = v_i^2 + 2g\Delta y = 0 + 2 \times 10 \times 100 = 2000$ . So W =  $KE_f - KE_i = \frac{1}{2}mv_f^2 = \frac{1}{2}1000 \times 2000 = 10^6$  J.

### 65 CM: Rigid Bodies

#### Answer: (D)

 $\tau = I\alpha \implies Mg^{\underline{L}}\sin\theta = \frac{1}{3}ML^2\alpha \implies \alpha = \frac{3}{2L}g\sin\theta$ . Here we used the moment of inertia around one end of the rod  $I = I_{CM} + Md^2 = \frac{1}{12}ML^2 + M(\frac{L}{2})^2 = \frac{1}{3}ML^2$ .

# 66 CM: Springs

### Answer: (C)

Two springs in series add like resistors in parallel, so  $k_{eq} = (\frac{1}{k_1} + \frac{1}{k_2})^{-1} = \frac{k}{2}$ . So,  $\omega = \sqrt{\frac{k_{eq}}{m}} = \sqrt{\frac{k}{2m}}$ .

# 67 CM: Hamiltonians

#### Answer: (E)

$$H = T + U = \frac{P^2}{2m} + \frac{1}{2}kx^2$$
.

### 68 EM: Gauss's Law

#### Answer: (A)

Using Gauss's law just inside every cavity in the conductor part, we see that  $E = 0 \implies q_{in} = 0$ , so the inner surface of each cavity develops a charge of minus what's held inside the cavity. So cavity A gets -2q, cavity B gets +4q and cavity C gets 0.

#### **EM: Induction** 69

Answer: (D) 
$$|\text{emf}| = \frac{N\partial\Phi}{\partial t} = \frac{N\Delta BA}{\Delta t} = \frac{250\times0.05\times0.2}{0.25} = 10\text{V}.$$

#### QM: Incidence on V 70

### Answer: (D)

If  $k_1 = k_2$ , then there should be no reflection since the potential is zero in that case. The only non trivial choice that satisfies this condition is (D).

#### 71 QM: Operators

We know that  $[\sigma_x, \sigma_y] = 2i\sigma_z \implies \sigma_x\sigma_y - \sigma_y\sigma_x = 2i\sigma_z$ . So  $\sigma_x\sigma_y$  probably won't be zero nor involving just  $\sigma_z$  without an i. You can also check by multiplying the matrices that  $\sigma_x\sigma_y=$  $-\sigma_y\sigma_x$ .

#### ST: Nuclear Physics 72

#### Answer: (C)

The famous nuclear binding energy per nucleon graph of all elements has a peak at <sup>56</sup>Fe.

#### 73 ST: Particle Physics

#### Answer: (E)

The muon, electron and tau particles are all leptons that have the same properties but just different masses.

#### 74 ST: Particle Physics

#### Answer: (A)

A lepton does not have any quark content. While baryons (like n and p) have three quarks each.

#### ST: Bragg's Law 75

#### Answer: (C)

Bragg's law is  $2d\sin\theta = m\lambda$ . So, we see that d (distance between layers) is in the order of  $\lambda$ . Then as KE =  $\frac{p^2}{2m}$  and using De Broglie's wavelength  $p = \frac{h}{\lambda} \implies$  KE =  $\frac{h^2}{2\lambda^2 m} = \frac{(6.6 \times 10^{-34})^2}{2 \times 9.1 \times 10^{-31} \times (0.4 \times 10^{-9})^2}$  J =  $\frac{(6.6 \times 10^{-34})^2}{2 \times 9.1 \times 10^{-31} \times (0.4 \times 10^{-9})^2 \times 1.6 \times 10^{-19}}$  eV  $\approx 10$  eV.

#### 76 EM: Magnetic Force

#### Answer: (E)

When the particle is undeflected, qvB = qE and hence  $v = \frac{E}{B}$ .

### 77 EM: Boundary Conditions

#### Answer: (C)

The boundary conditions for B as seen by Maxwell's equations ensure that the perpendicular component of B is continuous across boundaries. So the perpendicular component of B vanishes also outside and hence B is totally tangential.

### 78 EM: Radiation

#### Answer: (B)

An oscillating charge has a far Poynting vector in the  $\hat{r}$  direction (radially outward), so it's to the right here.

#### 79 CM: Forces

#### Answer: (C)

For  $m_2$ , Newton's second law is  $m_2g - T = m_2a \implies T = m_2(g - a) \implies T < m_2g$ .

#### 80 CM: Gravitation

#### Answer: (D)

Any of the two small masses feels a total radial force of  $\frac{GmM}{R^2} + \frac{Gmm}{(2R)^2} = \frac{Gm}{R^2}(M + \frac{m}{4})$ . So, they feel an effective big mass  $M' = M + \frac{m}{4}$ .

#### 81 WO: Thin Films

### Answer: (C)

There is a phase shift of  $\pi$  between the two rays reflected from the first and second face (because one of them is reflected by  $n_2 > n_1$  and the other is not). So the condition for constructive interference is  $(m + \frac{1}{2}) \frac{\lambda}{n} = 2t$ . Therefore,  $(m + \frac{1}{2}) \lambda = 1500$  nm. If m = 0,  $\lambda = 3000$  nm. If m = 1,  $\lambda = 1000$  nm. If m = 2,  $\lambda = 600$  nm. Hence 600 nm is one of the strongly reflected wavelengths. (Note that we divided the wavelength by n since when it travels inside the thin film, it has this reduced wavelength)

#### 82 WO: Refraction

#### Answer: (D)

The beam is reflected with an angle of  $60^{\circ}$ , so the angle of incidence inside is also  $60^{\circ}$ , and from the right angle triangle, we see that the angle of refraction inside is just  $90^{\circ}$  -  $60^{\circ} = 30^{\circ}$ . And hence our angle outside satisfies Snell's law  $1 \sin \alpha = n \sin 30^{\circ}$ . Therefore,  $\alpha = \sin^{-1} \frac{n}{2}$ .

#### 83 ST: Radiation Power

#### Answer: (C)

Power received = 3 (1000) = 3000 W. For the reflective surface, we have to double the power to look at the total force as the power is received then emitted. So,  $P = F v \implies 6000 = F \times 3 \times 10^8 \implies F = 2 \times 10^{-5} N$ .

### 84 SR: Doppler Shift

Answer: (C)

$$\frac{\lambda'}{\lambda} = \sqrt{\frac{1+\beta}{1-\beta}} \implies 9 = \frac{1+\beta}{1-\beta} \implies 9 - 9\beta = 1 + \beta \implies 10\beta = 8 \implies \beta = 0.8.$$

### 85 SR: Energy-Momentum

Answer: (A)

$$\beta = 0.6 \implies \gamma = \frac{1}{0.8} = \frac{5}{4}$$
. And  $W = E_f - E_i = \gamma mc^2 - mc^2 = (\gamma - 1)mc^2 = (\frac{5}{4} - 1)mc^2 = \frac{1}{4}mc^2$ .

### 86 ST: Hall Effect

Answer: (E)

Hall effect is always used to find the sign of charge carriers.

### 87 CM: Energy

Answer: (E)

Energy is conserved here, potential energy gets transferred to kinetic. Also, we can measure g by this setup, by measuring the time the mass drops and from that deduce a and from that calculate g. Momentum conservation, however, is not valid here since gravity acts like an external force that changes the momentum of the two masses.

### 88 CM: Relative Velocity

Answer: (B)

The velocity of the payload is equal to 100 m/s north + g  $\Delta t$  down = 100 m/s north + 40 m/s down. So relative to the plane, we just subtract the velocity of the plane which is 100 m/s north and we get 40 m/s down.

# 89 CM: Terminal Speed

Answer: (B)

Terminal speed happens when m g = drag force = c  $v^2$ . Hence,  $\frac{v_1^2}{v_2^2} = \frac{m_1}{m_2} \implies \frac{v_1}{v_2} = \sqrt{2}$ .

#### 90 CM: SHM

Answer: (A)

 $E_{tot} = \frac{1}{2}kA^2 = \frac{1}{2}mv_{max}^2 = \frac{1}{2}0.3 \times (0.04)^2 = 0.24$  mJ. Note that the maximum speed is the given 0.04 m/s since it occurs at the equilibrium position.

#### 91 TD: Carnot

Answer: (E)

The Carnot effeciency equation is  $\frac{W}{Q_H} = 1 - \frac{T_C}{T_H} = 1 - \frac{295}{305} = \frac{10}{305} \implies \frac{Q_H}{W} = \frac{305}{10} = 30.5.$ 

17

If we wanted to be exact and compare to  $Q_C$  instead, note that  $\frac{Q_H}{Q_C} = \frac{T_H}{T_C} = \frac{295}{305}$  and hence  $\frac{Q_C}{W} = \frac{295}{305} \frac{305}{10} = 29.5$ , so we are still fine.

### 92 QM: Particle in a Box

#### Answer: (B)

 $E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2}$ , so we are clearly in the n = 3 state, which has the probability density function as shown in Fig (1). So from symmetry, we see that the needed probability is  $\frac{1}{6}$ .

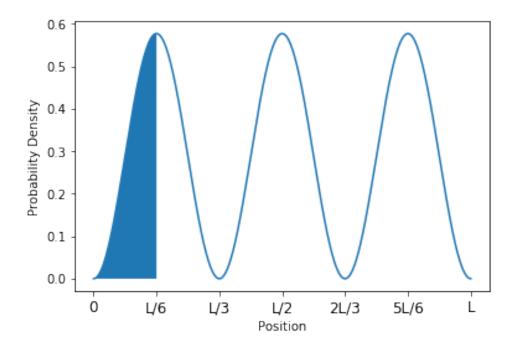


Figure 1:  $|\psi|^2$  for the n = 3 wavefunction for a particle in a box of length L

### 93 SR: 4-Vectors

#### Answer: (E)

 $E^2 - (pc)^2 = (mc^2)^2$  is an invariant.

### 94 EM: Circuits

#### Answer: (D)

The initial charge on the capacitor is  $Q_0 = CV_0$ . This charge is conserved, so in the final situation, it is still the total charge. We can combine the two empty capacitors in series to have a capacitance of  $\frac{C}{2}$ . Now the charge on this equivalent capacitor and the charge on the first capacitor must add up to the original charge  $Q_1 + Q_2 = Q_0 = CV_0$ . And also they share the same voltage  $V_1 = V_2 = V$ . So, the first equation becomes  $CV + \frac{C}{2}V = CV_0 \implies V = \frac{2}{3}V_0$ .

### 95 EM: Potential

#### Answer: (A)

In both cases, the integral for V is very simple V =  $\int \frac{kdq}{r} = \frac{kQ}{r}$  since r is constant for both cases. r at the center is the radius R, and r outside is just  $\sqrt{R^2 + x^2} = \sqrt{2}$ . So,  $\Delta U = q\Delta V = kqQ(\frac{1}{R} - \frac{1}{r}) = 9 \times 10^9 \times 5 \times 10^{-6} \times 3 \times 10^{-6} (1 - \frac{1}{\sqrt{2}}) = 135 \times 10^{-3} (1 - 0.707) \approx 39.5 \times 10^{-3} J$ .

#### 96 EM: Sources of B

#### Answer: (C)

Diamagnetic materials lower the magnetic field magnitude (B) but the direction remains the same.

### 97 CM: Lagrangians

#### Answer: (A)

L = T - U. Let's have a look at the gravitational potential here. Defining the reference for y = 0 at the top wall, we get  $U_g = -mg(h_1 + h_2) = -mga(\cos\theta_1 + \cos\theta_2)$ . If the angles are small, we can expand the cosine into  $\cos\theta \approx 1 - \frac{\theta^2}{2}$ . Therefore, for one of the two masses  $U_g = -mga(1 - \frac{\theta^2}{2}) = -mga + \frac{mga\theta^2}{2}$ . Now let's get rid of the constant potential that appears (-mga) by shifting the reference down at a distance a below the wall, and hence the gravititional potential becomes  $U_g = \frac{mga}{2}(\theta_1^2 + \theta_2^2)$ . Therefore, only choice (A) gives a form that is T - U.

### 98 ST: Bragg's Law

#### Answer: (D)

m  $\lambda = 2d\sin\theta$ . So, to have two angles, m must go up to 2. So,  $2\lambda = 2d\sin\theta \implies \lambda = d\sin\theta$ . Therefore,  $\lambda < d$ .

### 99 QM: Selection Rules

#### Answer: (B)

 $\Delta l = 0$  is a forbidden transition, so choice (B) is the only correct option. Only  $\Delta l = \pm 1$  are allowed.

### 100 ST: Mössbauer Spectroscopy

#### Answer: (C)

The velocity graph we see is a result of some tiny Doppler shifts, so from that we can deduce the shift in the wavelength, and hence the change in energy. And from the uncertainty principle, we can approximate the change in time, which is the lifetime needed.

So,  $v \approx 1$  mm/s means  $\beta = \frac{1}{3} \times 10^{-11}$ . Since this is a very tiny  $\beta$ , we use the usual approximation in Doppler to get,  $\frac{\lambda'}{\lambda} \approx (1+\beta)^2 \approx 1+2\beta$ . So, to first order,  $\frac{\Delta\lambda}{\lambda} = 2\beta$ .

Relating this to the energy shift, note that 
$$E = \frac{hc}{\lambda}$$
, so  $\Delta E = \frac{\partial E}{\partial \lambda} \Delta \lambda = \frac{hc}{\lambda^2} \Delta \lambda = \frac{E}{\lambda} \Delta \lambda$ . Hence,  $\frac{\Delta E}{E} = \frac{\Delta \lambda}{\lambda} = 2\beta \implies \Delta E = 2\beta E$ . And using  $\Delta E \Delta t \approx \hbar \implies \Delta t \approx \frac{6.6 \times 10^{-34}}{2\pi \times \frac{2}{3} \times 10^{-11} \times 14.4 \times 10^{3} \times 1.6 \times 10^{-19}} \approx 10^{-7} \text{s}.$